# Relativity of magnetic and electric fields <br> Masatsugu Sei Suzuki <br> Department of Physics, SUNY at Bimghamton <br> (Date: January 13, 2012) 

## 1. Lorentz transformation

### 1.1 Derivation of Lorentz transformation



We consider a Galilean transformation given by

$$
\begin{aligned}
& x^{\prime}=x-v t \\
& x=x^{\prime}+v t^{\prime} \\
& t^{\prime}=t \\
& \frac{d x^{\prime}}{d t^{\prime}}=\frac{d x}{d t^{\prime}}-v \frac{d t}{d t^{\prime}}=\frac{d x}{d t}-v \\
& u^{\prime}=u-v
\end{aligned}
$$

We know that the velocity of light remains unchanged under a transformation (so-called the Lorentz transformation) satisfying the principle of relativity. This implies that the Lorentz transformation is not the same as the Galilean transformation.

Here we assume that

$$
\begin{aligned}
& x^{\prime}=\gamma(x-v t) \\
& x=\gamma\left(x^{\prime}+v t^{\prime}\right)
\end{aligned}
$$

from the symmetry of transformation
What is the value of $\gamma$ ?
(i) The light is emitted at

$$
\begin{aligned}
& t=t^{\prime}=0 \\
& x=x^{\prime}=0
\end{aligned}
$$

(initially). The speed of light (in vacuum) is the same in all internal reference frames; it always has the value $c$.

$$
\frac{x^{\prime}}{t^{\prime}}=\frac{x}{t}=c
$$

## ((Mathematica))

Derivation of Lorentz transformation

$$
\begin{aligned}
& \text { eq1 }=x==\gamma\left(x^{\prime}+v t^{\prime}\right) ; e q 2=x^{\prime}==\gamma(x-v t) \\
& x^{\prime}=(-t v+x) \gamma \\
& \text { eq3 }=\text { Solve }\left[\{e q 1 \text {, eq2 }\},\left\{x^{\prime}, t^{\prime}\right\}\right] / / S i m p l i f y / / F l a t t e n \\
& \left\{t^{\prime} \rightarrow \frac{x\left(\frac{1}{\gamma}-\gamma\right)}{x^{\prime} v}+t \gamma, x^{\prime} \rightarrow(-t v+x) \gamma\right\} \\
& \text { eq4 }=\frac{x^{\prime}}{t^{\prime}}=\frac{x}{t} / . \text { eq3 // Simplify } \\
& \frac{v(-t v+x) \gamma^{2}}{x+t v \gamma^{2}-x \gamma^{2}}=\frac{x}{t} \\
& \text { eq5 }=\text { eq4/. }\{x \rightarrow c \text { t }\} / / \text { Simplify } \\
& -\frac{v(-C+v) \gamma^{2}}{c-c \gamma^{2}+v \gamma^{2}}=c \\
& \text { eq6=Solve [eq5, } \gamma \text { ] } \\
& \left\{\left\{\gamma \rightarrow-\frac{\text { i } C}{\sqrt{-C^{2}+V^{2}}}\right\},\left\{\gamma \rightarrow \frac{\text { i } C}{\sqrt{-\mathrm{C}^{2}+v^{2}}}\right\}\right\} \\
& \text { T'=t'/.eq3/.eq6[[2]]//Simplify } \\
& \frac{i 1\left(c^{2} t-v x\right)}{c \sqrt{-c^{2}+v^{2}}} \\
& x^{\prime} / . e^{3} \\
& \text { (-t v+x) } \gamma \\
& \text { X'=x'/.eq3/.eq6[[2]]//Simplify } \\
& \frac{\text { i } c(-t v+x)}{\sqrt{-c^{2}+v^{2}}}
\end{aligned}
$$

Then we have

$$
\begin{aligned}
& x^{\prime}=\gamma(x-v t)=\gamma(x-\beta c t) \\
& t^{\prime}=\gamma\left(t-\frac{\beta}{c} x\right)
\end{aligned}
$$

where

$$
\gamma=\frac{1}{\sqrt{1-\beta^{2}}}
$$

with

$$
\beta=\frac{v}{c}
$$

Note that $\gamma$ is expanded as

$$
\gamma=1+\frac{\beta^{2}}{2}+\frac{3 \beta^{4}}{8}+O\left(\beta^{6}\right)
$$

in the limit of $\beta \rightarrow 0$.
For convenience, we introduce

$$
x_{4}=i c t
$$

or

$$
-i \frac{x_{4}}{c}=t
$$

Then we have

$$
\begin{aligned}
& x_{1}^{\prime}=\gamma\left(x_{1}+i \beta x_{4}\right) \\
& x_{4}^{\prime}=\gamma\left(-i \beta x_{1}+x_{4}\right)
\end{aligned}
$$

or

$$
\left(\begin{array}{l}
x_{1}^{\prime} \\
x_{2}^{\prime} \\
x_{3}^{\prime} \\
x_{4}^{\prime}
\end{array}\right)=\left(\begin{array}{cccc}
\gamma & 0 & 0 & i \beta \gamma \\
0 & 1 & 0 & 0 \\
0 & 0 & 1 & 0 \\
-i \beta \gamma & 0 & 0 & \gamma
\end{array}\right)\left(\begin{array}{l}
x_{1} \\
x_{2} \\
x_{3} \\
x_{4}
\end{array}\right)
$$

or

$$
\begin{aligned}
& x^{\prime}=a x \\
& x_{1}=\gamma\left(x_{1}^{\prime}-i \beta x_{4}{ }^{\prime}\right) \\
& x_{4}=\gamma\left(i \beta x_{1}^{\prime}+x_{4}{ }^{\prime}\right)
\end{aligned}
$$

or in the matrix form,

$$
\begin{aligned}
& \left(\begin{array}{l}
x_{1} \\
x_{2} \\
x_{3} \\
x_{4}
\end{array}\right)=\left(\begin{array}{cccc}
\gamma & 0 & 0 & -i \beta \gamma \\
0 & 1 & 0 & 0 \\
0 & 0 & 1 & 0 \\
i \beta \gamma & 0 & 0 & \gamma
\end{array}\right)\left(\begin{array}{l}
x_{1}^{\prime} \\
x_{2}^{\prime} \\
x_{3}^{\prime} \\
x_{4}^{\prime}
\end{array}\right) \\
& x=a^{-1} x^{\prime}
\end{aligned}
$$

Note that $a^{-1}=a^{T}$

$$
x_{\mu}=\left(a^{-1}\right)_{\mu \nu}=\left(a^{T}\right)_{\mu \nu}=a_{v \mu \mu} x_{v}^{\prime}
$$

### 1.2 Lorentz contraction

Imagine a stick moving to the right at the velocity $v$. Its rest length (that is, its length measured in $S^{\prime}$ ) is $\Delta x_{1}{ }^{\prime}$.
We measure the distance of the stick under the condition that $\Delta x_{4}=0$. Since

$$
\Delta x_{1}^{\prime}=\gamma\left(\Delta x_{1}+i \beta \Delta x_{4}\right)=\gamma \Delta x_{1}
$$

or

$$
\Delta x_{1}=\frac{1}{\gamma} \Delta x_{1}^{\prime}=\sqrt{1-\beta^{2}} \Delta x_{1}^{\prime}=\sqrt{1-\beta^{2}} \Delta x_{0}
$$

The length of the stick measure in $\mathrm{S}\left(\Delta x_{1}\right)$ is shorter than that observed in $\mathrm{S}^{\prime}\left(\Delta x_{1}\right.$, proper length)

### 1.3 Time dilation

We are watching one moving clock moving to the right at the velocity $v$.

$$
\Delta x_{4}=\gamma\left(i \beta \Delta x_{1}{ }^{\prime}+\Delta x_{4}{ }^{\prime}\right)
$$

with $\Delta x_{1}{ }^{\prime}=0$. Then we have

$$
\Delta x_{4}=\gamma \Delta x_{4}^{\prime}>\Delta x_{4}{ }^{\prime} .
$$

or

$$
\Delta t=\frac{1}{\sqrt{1-\beta^{2}}} \Delta t^{\prime}=\frac{1}{\sqrt{1-\beta^{2}}} \Delta t_{0} \geq \Delta t_{0}
$$

The time in $\mathrm{S}(\Delta t)$ is longer than that observed in $\mathrm{S}^{\prime}\left(\Delta t_{0}\right.$, proper time). The moving clocks run slow

### 1.4 Proper time

$$
\left(d x_{\mu}^{\prime}\right)^{\prime 2}=a_{\mu \lambda} a_{\mu \sigma} d x_{\lambda} d x_{\sigma}=\delta_{\lambda \sigma} d x_{\lambda} d x_{\sigma}=\left(d x_{\mu}\right)^{2}
$$

We define the proper time as

$$
\begin{aligned}
& (d s)^{2}=c^{2}(d t)^{2}-\left(d x_{1}\right)^{2}-\left(d x_{2}\right)^{2}-\left(d x_{3}\right)^{2}=c^{2}\left(d t^{\prime}\right)^{2}-\left(d x_{1}\right)^{2}-\left(d x_{2}\right)^{2}-\left(d x_{3}\right)^{2} \\
& (d s)^{2}=c^{2}(d t)^{2}\left\{1-\frac{1}{c^{2}}\left[\left(\frac{d x_{1}}{d t}\right)^{2}+\left(\frac{d x_{2}}{d t}\right)^{2}+\left(\frac{d x_{3}}{d t}\right)^{2}\right]\right\}=c^{2}(d t)^{2}\left[1-\beta^{2}\right]
\end{aligned}
$$

or

$$
d \tau=\frac{d s}{c}=d t \sqrt{1-\beta^{2}}
$$

where $\tau$ is a proper time.

### 1.5 Notation of four vector

Four vector notation

$$
\begin{aligned}
& b=\left(\begin{array}{l}
b_{1} \\
b_{2} \\
b_{3} \\
b_{4}
\end{array}\right)=\left(\begin{array}{l}
b_{1} \\
b_{2} \\
b_{3} \\
i_{0}
\end{array}\right) \\
& b_{\mu}(\mu=1,2,3, \text { and } 4)
\end{aligned}
$$

where

$$
\begin{array}{ll}
b_{1}, b_{2}, b_{3}: & \text { real } \\
b_{4}=i b_{0} & \text { purely imaginary }
\end{array}
$$

((Note))
We use the Einstein convention, in which repeated indices are summed.

$$
\begin{aligned}
& i, j, k(=1-3) \\
& \mu, v, \lambda(=1,-4)
\end{aligned}
$$

The co-ordinate vector

$$
x=\left(\begin{array}{l}
x_{1} \\
x_{2} \\
x_{3} \\
x_{4}
\end{array}\right)=\left(\begin{array}{l}
x_{1} \\
x_{2} \\
x_{3} \\
i x_{0}
\end{array}\right)
$$

$$
b_{\mu}(\mu=1,2,3, \text { and } 4)
$$

Under a Lorentz transformation, we have

$$
x_{\mu}{ }^{\prime}=a_{\mu \nu} x_{v}
$$

instead of

$$
x_{\mu}^{\prime}=\sum_{v} a_{\mu \nu} x_{v}
$$

where

$$
\begin{aligned}
& a_{\mu \nu} a_{\mu \lambda}=\delta_{\nu \lambda} \\
& \left(a^{-1} a\right)_{\lambda \nu}=\left(a^{-1}\right)_{\lambda \mu} a_{\mu \nu}=\left(a^{T}\right)_{\lambda \mu} a_{\mu \nu}=a_{\mu \lambda} a_{\mu \nu}=\delta_{\lambda \nu}
\end{aligned}
$$

Note that

$$
a^{-1}=a^{T} \quad(\text { transpose matrix })
$$

(1)

$$
x_{\mu}^{\prime} x_{\mu}^{\prime}=a_{\mu \nu} x_{\nu} a_{\mu \lambda} x_{\lambda}=a_{\mu \nu} a_{\mu \lambda} x_{\nu} x_{\lambda}=\delta_{\mu \lambda} x_{\nu} x_{\lambda}=x_{\mu} x_{\mu}
$$

(2)

$$
\begin{aligned}
& x_{\mu}{ }^{\prime}=a_{\mu \lambda} x_{\lambda} \\
& a_{\mu \nu} x_{\mu}{ }^{\prime}=a_{\mu \nu} a_{\mu \lambda} x_{\lambda}=\delta_{\nu \lambda} x_{\lambda}=x_{v}
\end{aligned}
$$

or

$$
x_{v}=a_{\mu v} x_{\mu}{ }^{\prime} \quad \text { or } \quad x_{\mu}=a_{v \mu} x_{v}{ }^{\prime}
$$

A four vector, by definition, transforms in the same way as $x_{\mu}$ under the Lorentz transformation.

$$
\frac{\partial}{\partial x_{\mu}{ }^{\prime}}=\frac{\partial x_{v}}{\partial x_{\mu}{ }^{\prime}} \frac{\partial}{\partial x_{v}}=a_{\mu \nu} \frac{\partial}{\partial x_{v}}
$$

where

$$
x_{\nu}=a_{\mu \nu} x_{\mu}{ }^{\prime}
$$

The scalar product $b \cdot c$ is defined by

$$
b \cdot c=b_{\mu} c_{\mu}
$$

It is invariant under the Lorentz transformation

$$
b_{\mu}^{\prime} c_{\mu}^{\prime}=a_{\mu \nu} b_{\nu} a_{\mu \lambda} b_{\lambda}=a_{\mu \nu} a_{\mu \lambda} a_{v} b_{\lambda}=\delta_{\nu \lambda} a_{v} b_{\lambda}=a \cdot b
$$

### 1.6 Four dimensional Laplacian operator

$$
\frac{\partial}{\partial x_{\mu}} \frac{\partial}{\partial x_{\mu}}=\frac{\partial^{2}}{\partial x_{1}{ }^{2}}+\frac{\partial^{2}}{\partial x_{2}{ }^{2}}+\frac{\partial^{2}}{\partial x_{3}{ }^{2}}+\frac{\partial^{2}}{\partial x_{4}{ }^{2}}=\nabla^{2}-\frac{1}{c^{2}} \frac{\partial^{2}}{\partial t^{2}}
$$

is invariant under the Lorentz transformation: Lorentz scalar

$$
\frac{\partial}{\partial x_{\mu}{ }^{\prime}} \frac{\partial}{\partial x_{\mu}{ }^{\prime}}=a_{\mu \nu} \frac{\partial}{\partial x_{v}} a_{\mu \mu} \frac{\partial}{\partial x_{\lambda}}=a_{\mu \nu} a_{\mu \lambda} \frac{\partial}{\partial x_{v}} \frac{\partial}{\partial x_{\lambda}}=\delta_{v \lambda} \frac{\partial}{\partial x_{v}} \frac{\partial}{\partial x_{\lambda}}=\frac{\partial}{\partial x_{v}} \frac{\partial}{\partial x_{v}}=\frac{\partial}{\partial x_{\mu}} \frac{\partial}{\partial x_{\mu}}
$$

### 1.7 A tensor of second rank

A tensor of second rank, $t_{\mu \nu}$, transforms as

$$
t_{\mu \nu}{ }^{\prime}=a_{\mu \lambda} a_{\nu \sigma} t_{\lambda \sigma}
$$

### 1.8 A tensor of third rank

A tensor of third rank, $t_{\mu \nu \lambda}$, transforms as

$$
t_{\mu \nu \lambda}{ }^{\prime}=a_{\mu \sigma} a_{\nu \rho} a_{\lambda \tau} t_{\sigma \rho \tau}
$$

((Note))
We make no distinction between a covariant and contravariant vector. We do not define the metric tensor $g_{\mu \nu}$.

## 2. Velocity, acceleration, and force <br> 2.1 Lorentz velocity transformation

$$
x_{\mu}=\left(a^{-1}\right)_{\mu \nu} x_{\nu}^{\prime}
$$

So we have

$$
\begin{array}{ll}
\left(a^{-1}\right)_{\mu \nu}=a_{v \mu}=\left(a^{T}\right)_{\mu \nu} & \\
x_{1}^{\prime}=\gamma\left(x_{1}-v t\right) & x_{1}=\gamma\left(x_{1}{ }^{\prime}+v t^{\prime}\right) \\
x_{2}^{\prime}=x_{2} & x_{2}=x_{2}{ }^{\prime} \\
x_{3}^{\prime}=x_{3} & x_{3}=x_{3}{ }^{\prime} \\
t^{\prime}=\gamma\left(t-\frac{\beta}{c} x_{1}\right) & t=\gamma\left(t^{\prime}+\frac{\beta}{c} x_{1}{ }^{\prime}\right)
\end{array}
$$

Suppose that an object has velocity components as measured in $S^{\prime}$ and $S$.

$$
\begin{array}{ll}
u_{1}^{\prime}=\frac{d x_{1}^{\prime}}{d t^{\prime}}=\frac{u_{1}-v}{1-\frac{\beta}{c} u_{1}} & u_{1}=\frac{d x_{1}}{d t^{\prime}}=\frac{u_{1}^{\prime}+v}{1+\frac{\beta}{c} u_{1}^{\prime}} \\
u_{2}^{\prime}=\frac{d x_{2}^{\prime}}{d t^{\prime}}=\frac{1}{\gamma} \frac{u_{2}}{1-\frac{\beta}{c} u_{1}} & u_{2}=\frac{d x_{2}}{d t}=\frac{1}{\gamma} \frac{u_{2}^{\prime}}{1+\frac{\beta}{c} u_{1}^{\prime}} \\
u_{3}^{\prime}=\frac{d x_{3}^{\prime}}{d t^{\prime}}=\frac{1}{\gamma} \frac{u_{3}}{1-\frac{\beta}{c} u_{1}} & u_{3}=\frac{d x_{3}}{d t}=\frac{1}{\gamma} \frac{u_{3}^{\prime}}{1+\frac{\beta}{c} u_{1}^{\prime}}
\end{array}
$$

The Lorentz transformation of a velocity less than $c$ never leads to a velocity greater than $c$. The relations reduce to the Galilean transformation for $v \ll c$.

Suppose that the particle is sphoton, and $u_{1}=c$ in the frame S. Then we have

$$
u_{1}^{\prime}=\frac{u_{1}-v}{1-\frac{\beta}{c} u_{1}}=\frac{c-v}{1-\frac{\beta}{c} c}=\frac{c-v}{1-\frac{v}{c}}=c
$$

((Example))

$$
\begin{array}{ll}
u_{1}^{\prime}=\frac{9}{10} c \\
v=\frac{9}{10} c & u_{1}=\frac{u_{1}^{\prime}+v}{1+\frac{\beta}{c} u_{1}^{\prime}}=\frac{\frac{9}{10} c+\frac{9}{10} c}{1+\frac{81}{100}}=\frac{180}{181} c<c
\end{array}
$$

whereas the Galilean transformation would have given

$$
u_{1}=u_{1}^{\prime}+v=\frac{9}{10} c+\frac{9}{10} c=\frac{9}{5} c>c
$$

### 2.2 Lorentz acceleration transformation

Similarly we have the acceleration components as measured in $S^{\prime}$ and $S$.

$$
\begin{aligned}
& a_{1}=\frac{d u_{1}}{d t}=\frac{d t^{\prime}}{d t} \frac{d}{d t^{\prime}}\left(\frac{u_{1}{ }^{\prime}+v}{1+\frac{\beta}{c} u_{1}{ }^{\prime}}\right)=\frac{1}{\gamma} \frac{1}{\left(1+\frac{\beta}{c} u_{1}{ }^{\prime}\right)} \frac{d}{d t^{\prime}}\left(\frac{u_{1}{ }^{\prime}+v}{1+\frac{\beta}{c} u_{1}{ }^{\prime}}\right)=\frac{1}{\gamma} \frac{\left(1-\beta^{2}\right) a_{1}{ }^{\prime}}{\left(1+\frac{\beta}{c} u_{1}{ }^{\prime}\right)^{3}}=\frac{1}{\gamma^{3}} \frac{a_{1}{ }^{\prime}}{\left(1+\frac{\beta}{c} u_{1}{ }^{\prime}\right)^{3}} \\
& a_{2}=\frac{d u_{2}}{d t}=\frac{1}{\gamma} \frac{d t^{\prime}}{d t} \frac{d}{d t^{\prime}}\left(\frac{u_{2}{ }^{\prime}}{1+\frac{\beta}{c} u_{1}{ }^{\prime}}\right)=\frac{1}{\gamma^{2}} \frac{1}{\left(1+\frac{\beta}{c} u_{1}{ }^{\prime}\right)} \frac{d}{d t^{\prime}}\left(\frac{u_{2}{ }^{\prime}}{1+\frac{\beta}{c} u_{1}^{\prime}}\right)=\frac{1}{\gamma^{2}} \frac{a_{2}{ }^{\prime}}{\left(1+\frac{\beta}{c} u_{1}{ }^{\prime}\right)^{2}}-\frac{1}{\gamma^{2}} \frac{a_{1}{ }^{\prime} a_{2} \frac{\beta}{\left(1+\frac{\beta}{c} u_{1}{ }^{\prime}\right)^{3}}}{a_{3}=\frac{d u_{3}}{d t}=\frac{1}{\gamma} \frac{d t^{\prime}}{d t} \frac{d}{d t^{\prime}}\left(\frac{u_{3}^{\prime}}{1+\frac{\beta}{c} u_{1}^{\prime}}\right)=\frac{1}{\gamma^{2}} \frac{1}{\left(1+\frac{\beta}{c} u_{1}{ }^{\prime}\right)} \frac{d}{d t^{\prime}}\left(\frac{u_{3}^{\prime}}{1+\frac{\beta}{c} u_{1}{ }^{\prime}}\right)=\frac{1}{\gamma^{2}} \frac{a_{3}^{\prime}}{\left.\left(1+\frac{\beta}{c} u_{1}\right)^{3}\right)^{3}}-\frac{1}{\gamma^{2}} \frac{a_{1}{ }^{\prime} a_{3} \frac{\beta}{c}}{\left(1+\frac{\beta}{c} u_{1}\right)^{3}}}
\end{aligned}
$$

where

$$
\frac{d t^{\prime}}{d t}=\frac{1}{\gamma} \frac{1}{\left(1+\frac{\beta}{c} u_{1}{ }^{\prime}\right)}
$$

The acceleration is a quantity of limited and questionable value in special relativity. Not only is it not an invariant, but the expressions for it are in general cumbersome, and moreover its different components transform in different ways.

### 2.3 Force F under the Lorentz transformation

Lorentz transformation

$$
\begin{aligned}
& x^{\prime}=\gamma(x-v t) \\
& y^{\prime}=y \\
& z^{\prime}=z \\
& t^{\prime}=\gamma\left(-\frac{v}{c^{2}} x_{1}+t\right)
\end{aligned}
$$

or

$$
\begin{aligned}
& x=\gamma\left(x^{\prime}+v t^{\prime}\right) \\
& y=y^{\prime} \\
& z=z^{\prime} \\
& t=\gamma\left(\frac{v}{c^{2}} x_{1}{ }^{\prime}+t^{\prime}\right) \\
& \frac{d p_{1}{ }^{\prime}}{d t^{\prime}}=\frac{\gamma\left(d p_{1}-\frac{\beta}{c} d E\right)}{\gamma\left(-\frac{v}{c^{2}} d x_{1}+d t\right)}=\frac{\gamma\left(\frac{d p_{1}}{d t}-\frac{\beta}{c} \frac{d E}{d t}\right)}{\gamma\left(-\frac{v}{c^{2}} \frac{d x_{1}}{d t}+1\right)}=\frac{\frac{d p_{1}}{d t}-\frac{\beta}{c} \frac{d E}{d t}}{1-\frac{v}{c^{2}} \frac{d x_{1}}{d t}}=\frac{F_{1}-\frac{\beta}{c} \frac{d E}{d t}}{1-\frac{v}{c^{2}} u_{1}}
\end{aligned}
$$

Since $\frac{d E}{d t}=\mathbf{F} \cdot \mathbf{u}$

$$
F_{1}{ }^{\prime}=\frac{d p_{1}{ }^{\prime}}{d t^{\prime}}=\frac{F_{1}-\frac{\beta}{c}(\mathbf{u} \cdot \mathbf{F})}{1-\frac{v}{c^{2}} u_{1}}
$$

Similarly

$$
\begin{aligned}
& F_{2}{ }^{\prime}=\frac{d p_{2}{ }^{\prime}}{d t^{\prime}}=\frac{d p_{2}}{\gamma\left(-\frac{v}{c^{2}} d x_{1}+d t\right)}=\frac{\frac{d p_{2}}{d t}}{\gamma\left(-\frac{v}{c^{2}} \frac{d x_{1}}{d t}+1\right)}=\frac{F_{2}}{\gamma\left(1-\frac{v}{c^{2}} u_{1}\right)} \\
& F_{3}{ }^{\prime}=\frac{d p_{3}{ }^{\prime}}{d t^{\prime}}=\frac{d p_{3}}{\gamma\left(-\frac{v}{c^{2}} d x_{1}+d t\right)}=\frac{\frac{d p_{3}}{d t}}{\gamma\left(-\frac{v}{c^{2}} \frac{d x_{1}}{d t}+1\right)}=\frac{F_{3}}{\gamma\left(1-\frac{v}{c^{2}} u_{1}\right)}
\end{aligned}
$$

We consider one special case when the particle is instantaneously at rest in $S$. So that $\boldsymbol{u}=0$.

$$
\begin{aligned}
& F_{1}^{\prime}=F_{1} \\
& F_{2}=\frac{F_{2}}{\gamma} \\
& F_{3}=\frac{F_{3}}{\gamma}
\end{aligned}
$$

The component of $\boldsymbol{F}$ parallel to the motion of $S^{\prime}$ is unchanged, whereas the components perpendicular are divided by $\gamma$.

## 3. Charge and current density

### 3.1 Charge density



We measure the distance of the cylinder under the condition that $\Delta x_{4}=0$. Since

$$
\Delta x_{1}^{\prime}=\gamma\left(\Delta x_{1}+i \beta \Delta x_{4}\right)=\gamma \Delta x_{1}
$$

or

$$
\Delta x_{1}=\frac{1}{\gamma} \Delta x_{1}^{\prime}=\sqrt{1-\beta^{2}} \Delta x_{1}^{\prime}
$$

we have

$$
L=\frac{1}{\gamma} L^{\prime}=\sqrt{1-\beta^{2}} L^{\prime}
$$

but with the same area $A$ (since dimension transverse to the motion are unchangeable. If we call $\rho^{\prime}\left(=\rho_{0}\right)$ the density of charges in the $S^{\prime}$ frame in which charges momentarily at rest, the total charge $Q$ is the same in any system,

$$
Q=\rho^{\prime} L^{\prime} A^{\prime}=\rho_{0} L^{\prime} A^{\prime}=\rho L A
$$

or

$$
\rho_{0} L^{\prime}=\rho L
$$

or

$$
\rho=\rho_{0} \frac{L^{\prime}}{L}=\gamma \rho_{0}
$$

### 3.2 Current density $J_{\mu}$

The current density $J_{\mu}$ is defined as

$$
J_{\mu}=(\mathbf{J}, i c \rho)=(\rho \mathbf{u}, i c \rho)
$$

where $\boldsymbol{u}$ is the velocity of the particle in the S frame. Evidently the charge density and current density go together to make a 4 vector.

$$
\begin{aligned}
& J_{\mu}^{\prime}=a_{\mu \nu} J_{v} \\
& J_{1}^{\prime}=\gamma\left(J_{1}+i \beta J_{4}\right)=\gamma\left(J_{1}-v \rho\right) \\
& \rho^{\prime}=\gamma\left(-\frac{\beta}{c} J_{1}+\rho\right)
\end{aligned}
$$

or

$$
\begin{aligned}
& J_{\mu}=a_{v \mu} J_{v}{ }^{\prime} \\
& J_{1}=\gamma\left(J_{1}{ }^{\prime}-i \beta J_{4}{ }^{\prime}\right)=\gamma\left(J_{1}{ }^{\prime}+\nu \rho^{\prime}\right) \\
& \rho=\gamma\left(\frac{\beta}{c} J_{1}{ }^{\prime}+\rho^{\prime}\right)
\end{aligned}
$$

Then we have

$$
\rho=\gamma\left(\frac{\beta}{c} J_{1}{ }^{\prime}+\rho^{\prime}\right)
$$

Note that

$$
\rho=\gamma \rho^{\prime}
$$

when $J_{1}{ }^{\prime}=0$.

### 3.3 Invariance under the Lorentz transformation

We know that $J_{\mu} J_{\mu}$ is invariant under the Lorentz transformation

$$
J^{\prime}{ }_{\mu} J_{\mu}^{\prime}=a_{\mu \nu} J_{V} a_{\mu \lambda} J_{\lambda}=a_{\mu \nu} a_{\mu \lambda} J_{\nu} J_{\lambda}=\delta_{\mu \lambda} J_{V} J_{\lambda}=J_{\mu} J_{\mu}
$$

or

$$
J_{\mu} J_{\mu}=\mathbf{J}^{2}-c^{2} \rho^{2}=\mathbf{J}^{\prime 2}-c^{2} \rho^{\prime 2}
$$

Suppose that $J^{\prime}=0\left(\right.$ or $\left.\boldsymbol{u}^{\prime}=0\right)$ in the $S^{\prime}$ frame, where the point charge is at rest. $J=\rho \mathbf{u}=\rho \mathbf{v}$ (the frame $S^{\prime}$ moves at the velocity $v$ relative to the frame $S$ ). Then we have

$$
\rho^{2} v^{2}-c^{2} \rho^{2}=0-c^{2} \rho^{\prime 2}=-c^{2} \rho_{0}^{2}
$$

or

$$
\rho \sqrt{1-\frac{v^{2}}{c^{2}}}=\rho_{0}, \quad \text { or } \quad \rho=\frac{\rho_{0}}{\sqrt{1-\frac{v^{2}}{c^{2}}}}=\gamma \rho_{0}
$$

## 4 Maxwell's equation field fensor

### 4.1 Four vectors for the vector potential and scalar potential

$$
\begin{aligned}
J_{\mu} & =(\mathbf{J}, i c \rho) \\
A_{\mu} & =\left(\mathbf{A}, i \frac{1}{c} \phi\right) \\
\partial_{\mu} & =\left(\nabla, \frac{\partial}{\partial(i c t)}\right)
\end{aligned}
$$

The equation of continuity;

$$
\partial_{\mu} J_{\mu}=\nabla \cdot \mathbf{J}+\left(-i \frac{1}{c} \frac{\partial}{\partial t}\right)(i c \rho)=\nabla \cdot \mathbf{J}+\frac{\partial}{\partial t} \rho=0
$$

Maxwell's equation;

$$
\begin{aligned}
& \nabla \cdot \mathbf{D}=\rho \\
& \nabla \cdot \mathbf{B}=0 \\
& \nabla \times \mathbf{E}=-\frac{\partial \mathbf{B}}{\partial t} \\
& \nabla \times \mathbf{H}=\mathbf{J}+\frac{\partial \mathbf{D}}{\partial t}
\end{aligned}
$$

where

$$
\begin{aligned}
& \mathbf{D}=\varepsilon_{0} \mathbf{E}+\mathbf{P} \\
& \mathbf{B}=\mu_{0}(\mathbf{H}+\mathbf{M})
\end{aligned}
$$

$$
\begin{aligned}
& \mathbf{E}=-\frac{\partial}{\partial t} \mathbf{A}-\nabla \phi \\
& \mathbf{B}=\nabla \times \mathbf{A}
\end{aligned}
$$

### 4.2. Gauge transformation

$$
\begin{aligned}
& A_{\mu}=\left(A, i \frac{1}{c} \phi\right) \\
& \mathbf{A}^{\prime}=\mathbf{A}+\nabla \lambda, \\
& \phi^{\prime}=\phi-\frac{\partial \lambda}{\partial}, \\
& A_{\mu}^{\prime}=A_{\mu}+\partial_{\mu} \lambda
\end{aligned}
$$

((Note))

$$
\begin{aligned}
& i \frac{1}{c} \phi^{\prime}=i \frac{1}{c} \phi+\frac{\partial \lambda}{\partial(i c t)} \\
& A_{4}^{\prime}=A_{4}+\frac{\partial \lambda}{\partial x_{4}}
\end{aligned}
$$

Lorentz gauge:

$$
\frac{\partial A_{\mu}}{\partial x_{\mu}}=\partial_{\mu} A_{\mu}=\nabla \cdot \mathbf{A}+\frac{\partial}{\partial(i c t)} i \frac{\phi}{c}=\nabla \cdot \mathbf{A}+\frac{1}{c^{2}} \frac{\partial \phi}{\partial t}=0
$$

### 4.3 Electromagnetic field tensor $F$

We define the field tensor as

$$
F_{\mu \nu}=\frac{\partial A_{v}}{\partial x_{\mu}}-\frac{\partial A_{\mu}}{\partial x_{v}}
$$

This tensor satisfies the Jacobi identity;

$$
\frac{\partial F_{\mu v}}{\partial x_{\lambda}}+\frac{\partial F_{\nu \lambda}}{\partial x_{\mu}}+\frac{\partial F_{\lambda \mu}}{\partial x_{v}}=0
$$

This equation holds automatically for the antisymmetric tensor
The magnetic field;

$$
\begin{aligned}
& F_{12}=\frac{\partial A_{2}}{\partial x_{1}}-\frac{\partial A_{1}}{\partial x_{2}}=B_{3} \\
& F_{23}=\frac{\partial A_{3}}{\partial x_{2}}-\frac{\partial A_{2}}{\partial x_{3}}=B_{1} \\
& F_{31}=\frac{\partial A_{1}}{\partial x_{3}}-\frac{\partial A_{3}}{\partial x_{1}}=B_{2}
\end{aligned}
$$

The electric field;

$$
\begin{aligned}
& F_{14}=\frac{\partial A_{4}}{\partial x_{1}}-\frac{\partial A_{1}}{\partial x_{4}}=\frac{E_{1}}{i c}=-\frac{i}{c} E_{1} \\
& F_{24}=\frac{\partial A_{4}}{\partial x_{2}}-\frac{\partial A_{2}}{\partial x_{4}}=\frac{E_{2}}{i c}=-\frac{i}{C} E_{2} \\
& F_{34}=\frac{\partial A_{4}}{\partial x_{3}}-\frac{\partial A_{3}}{\partial x_{4}}=\frac{E_{3}}{i c}=-\frac{i}{C} E_{3}
\end{aligned}
$$

The field tensor is an anti-symmetric tensor of second rank and hence, has 6 independent components.

Electromagnetic field tensor;

$$
F_{\mu \nu}=\left(\begin{array}{cccc}
0 & B_{3} & -B_{2} & -\frac{i}{c} E_{1} \\
-B_{3} & 0 & B_{1} & -\frac{i}{c} E_{2} \\
B_{2} & -B_{1} & 0 & -\frac{i}{c} E_{3} \\
\frac{i}{c} E_{1} & \frac{i}{c} E_{2} & \frac{i}{c} E_{3} & 0
\end{array}\right)
$$

We show that

$$
\begin{aligned}
& F_{\mu \nu}{ }^{\prime}=\frac{\partial A_{v}{ }^{\prime}}{\partial x_{\mu}{ }^{\prime}}-\frac{\partial A_{\mu}{ }^{\prime}}{\partial x_{\nu}{ }^{\prime}}=a_{\mu \sigma} a_{v \tau} F_{\sigma \tau} \\
& \frac{\partial}{\partial x_{\mu}{ }^{\prime}}=a_{\mu \nu} \frac{\partial}{\partial x_{\nu}}, \quad \text { and } \quad A_{\mu}{ }^{\prime}=a_{\mu \lambda} A_{\lambda}
\end{aligned}
$$

$$
F_{\mu \nu}{ }^{\prime}=\frac{\partial A_{v}{ }^{\prime}}{\partial x_{\mu}{ }^{\prime}}-\frac{\partial A_{\mu}{ }^{\prime}}{\partial x_{v}{ }^{\prime}}=a_{\mu \sigma} \frac{\partial A_{v}{ }^{\prime}}{\partial x_{\sigma}}-a_{v \tau} \frac{\partial A_{\mu}{ }^{\prime}}{\partial x_{\tau}}=a_{\mu \sigma} \frac{\partial\left(a_{v \tau} A_{\tau}\right)}{\partial x_{\sigma}}-a_{v \tau} \frac{\partial\left(a_{\mu \sigma} A_{\sigma}\right)}{\partial x_{\tau}}
$$

or

$$
F_{\mu \nu}^{\prime}=a_{\mu \sigma} a_{v \tau}\left(\frac{\partial A_{\tau}}{\partial x_{\sigma}}-\frac{\partial A_{\sigma}}{\partial x_{\tau}}\right)=a_{\mu \sigma} a_{v \tau} F_{\sigma \tau}
$$

### 4.4 Maxwell's equation (1)

The Maxwell's equation is given by

$$
\begin{aligned}
& \frac{\partial F_{\mu \nu}}{\partial x_{v}}=\mu_{0} J_{\mu} \\
& \frac{\partial F_{1 \mu}}{\partial x_{\mu}}=\frac{\partial F_{11}}{\partial x_{1}}+\frac{\partial F_{12}}{\partial x_{2}}+\frac{\partial F_{13}}{\partial x_{3}}+\frac{\partial F_{14}}{\partial x_{4}}=\frac{\partial B_{3}}{\partial x_{2}}-\frac{\partial B_{2}}{\partial x_{3}}-\frac{\frac{i}{c}}{i c} \frac{\partial E_{1}}{\partial t}=(\nabla \times \mathbf{B})_{1}-\frac{1}{c^{2}} \frac{\partial E_{1}}{\partial t} \\
& (\nabla \times \mathbf{B})_{1}=\varepsilon_{0} \mu_{0} \frac{\partial E_{1}}{\partial t}+\mu_{0} J_{1}=\mu_{0}\left(J_{1}+\varepsilon_{0} E_{1}\right) \\
& \frac{\partial F_{2 \mu}}{\partial x_{\mu}}=\frac{\partial F_{21}}{\partial x_{1}}+\frac{\partial F_{22}}{\partial x_{2}}+\frac{\partial F_{23}}{\partial x_{3}}+\frac{\partial F_{24}}{\partial x_{4}}=-\frac{\partial B_{3}}{\partial x_{1}}+\frac{\partial B_{1}}{\partial x_{3}}-\frac{\frac{i}{c}}{i c} \frac{\partial E_{2}}{\partial t}=(\nabla \times \mathbf{B})_{2}-\frac{1}{c^{2}} \frac{\partial E_{2}}{\partial t} \\
& (\nabla \times \mathbf{B})_{2}=\varepsilon_{0} \mu_{0} \frac{\partial E_{2}}{\partial t}+\mu_{0} J_{2}=\mu_{0}\left(J_{2}+\varepsilon_{0} E_{2}\right) \\
& \frac{\partial F_{3 \mu}}{\partial x_{\mu}}=\frac{\partial F_{31}}{\partial x_{1}}+\frac{\partial F_{32}}{\partial x_{2}}+\frac{\partial F_{33}}{\partial x_{3}}+\frac{\partial F_{34}}{\partial x_{4}}=\frac{\partial B_{2}}{\partial x_{1}}-\frac{\partial B_{1}}{\partial x_{2}}-\frac{i}{c} \frac{\partial E_{3}}{\partial t}=(\nabla \times \mathbf{B})_{3}-\frac{1}{c^{2}} \frac{\partial E_{3}}{\partial t} \\
& (\nabla \times \mathbf{B})_{3}=\varepsilon_{0} \mu_{0} \frac{\partial E_{3}}{\partial t}+\mu_{0} J_{3}=\mu_{0}\left(J_{3}+\varepsilon_{0} E_{3}\right)
\end{aligned}
$$

$$
\begin{aligned}
& \frac{\partial F_{4 \mu}}{\partial x_{\mu}}=\frac{\partial F_{41}}{\partial x_{1}}+\frac{\partial F_{42}}{\partial x_{2}}+\frac{\partial F_{43}}{\partial x_{3}}+\frac{\partial F_{44}}{\partial x_{4}}=\frac{i}{c} \frac{\partial E_{1}}{\partial x_{1}}+\frac{i}{c} \frac{\partial E_{2}}{\partial x_{2}}+\frac{i}{c} \frac{\partial E_{3}}{\partial x_{3}}=\frac{i}{c} \nabla \cdot \mathbf{E}=\mu_{0} J_{4} \\
& \frac{i}{c} \nabla \cdot \mathbf{E}=\mu_{0} i c \rho \\
& \nabla \cdot \mathbf{E}=\mu_{0} c^{2} \rho=\frac{\rho}{\varepsilon_{0}}
\end{aligned}
$$

((Note))

$$
\begin{aligned}
& \mathbf{B}=\nabla \times \mathbf{A}=\left|\begin{array}{ccc}
\hat{x}_{1} & \hat{x}_{2} & \hat{x}_{3} \\
\frac{\partial}{\partial x_{1}} & \frac{\partial}{\partial x_{2}} & \frac{\partial}{\partial x_{3}} \\
A_{1} & A_{2} & A_{3}
\end{array}\right|=\left(\frac{\partial A_{3}}{\partial x_{2}}-\frac{\partial A_{2}}{\partial x_{3}}, \frac{\partial A_{1}}{\partial x_{3}}-\frac{\partial A_{3}}{\partial x_{1}}, \frac{\partial A_{2}}{\partial x_{1}}-\frac{\partial A_{1}}{\partial x_{2}}\right) \\
& \mathbf{E}=-\nabla \phi-\frac{\partial \mathbf{A}}{\partial t}=\left(-\frac{\partial \phi}{\partial x_{1}}-i c \frac{\partial A_{1}}{\partial x_{4}},-\frac{\partial \phi}{\partial x_{2}}-i c \frac{\partial A_{2}}{\partial x_{4}},-\frac{\partial \phi}{\partial x_{3}}-i c \frac{\partial A_{3}}{\partial x_{4}}\right)
\end{aligned}
$$

or

$$
\mathbf{E}=\operatorname{ic}\left(\frac{\partial A_{4}}{\partial x_{1}}-\frac{\partial A_{1}}{\partial x_{4}}, \frac{\partial A_{4}}{\partial x_{2}}-\frac{\partial A_{2}}{\partial x_{4}}, \frac{\partial A_{4}}{\partial x_{3}}-\frac{\partial A_{3}}{\partial x_{4}}\right)
$$

where

$$
\phi=\frac{c}{i} A_{4} .
$$

### 4.5 Invariants of the field

$F_{\mu \nu} F_{\mu \nu}$ is invariant under the Lorentz transformation

$$
\begin{aligned}
& F_{\mu \nu}^{\prime} F_{\mu \nu}^{\prime}=a_{\mu \lambda} a_{\nu \rho} a_{\mu \sigma} a_{\nu \tau} F_{\lambda \rho} F_{\sigma \tau}=\delta_{\lambda \sigma} \delta_{\rho \tau} F_{\lambda \rho} F_{\sigma \tau}=F_{\lambda \rho} F_{\lambda \rho}=F_{\mu \nu} F_{\mu \nu} \\
& F_{\mu \nu} F_{\mu \nu}=2\left[B_{1}^{2}+B_{2}^{2}+B_{3}^{2}-\frac{1}{c^{2}}\left(E_{1}^{2}+E_{2}^{2}+E_{3}^{2}\right)\right]=\text { in variant }
\end{aligned}
$$

A further invariant is obtained by contraction of the field tensor with the "completely anti-symmetric unit tensor of fourth rank" defined by

$$
\varepsilon_{\kappa \lambda \mu \nu}=
$$

0 if two indices are equal,

1 if ( $\kappa \lambda \mu \nu$ ) is an even permutation of (1234), and -1 if $(\kappa \lambda \mu \nu)$ is an odd permutation of (1234).
(Levi-Civita tensor)
One may be convinced easily that $\varepsilon_{\kappa \lambda \mu \nu}$ is a tensor of rank 4 because

$$
\varepsilon_{\kappa^{\prime} \lambda^{\prime} \mu^{\prime} \nu^{\prime}}=a_{\kappa^{\prime} k} a_{\lambda^{\prime} \lambda} a_{\mu^{\prime}, \mu} a_{v^{\prime} \nu} \varepsilon_{\kappa \lambda \mu \nu \nu}
$$

Now we consider

$$
\varepsilon_{\kappa \lambda \lambda \nu} F_{\kappa \lambda} F_{\mu \nu}=\varepsilon_{1234} F_{12} F_{34}+\varepsilon_{1324} F_{13} F_{24}+\ldots=-\frac{8 i}{c} \mathbf{E} \cdot \mathbf{B}
$$

So the scalar product $\mathbf{E} \cdot \mathbf{B}$ is Lorentz invariant,

### 4.6 Equation of continuity

$$
\begin{aligned}
& F_{\mu \nu}=-F_{v \mu} \\
& F_{\mu \nu}=\frac{F_{\mu \nu}+F_{\mu \nu}}{2}=\frac{F_{\mu \nu}-F_{v \mu}}{2} \\
& \frac{\partial}{\partial x_{\mu}}\left(\frac{\partial F_{\mu \nu}}{\partial x_{v}}\right)=\frac{1}{2} \frac{\partial}{\partial x_{\mu}} \frac{\partial}{\partial x_{v}}\left(F_{\mu \nu}-F_{\nu \mu}\right)=\frac{1}{2}\left(\frac{\partial}{\partial x_{\mu}} \frac{\partial}{\partial x_{\nu}} F_{\mu \nu}-\frac{\partial}{\partial x_{v}} \frac{\partial}{\partial x_{\mu}} F_{\nu \mu}\right)=0
\end{aligned}
$$

Since

$$
\frac{\partial F_{\mu v}}{\partial x_{v}}=\mu_{0} J_{\mu}
$$

we have

$$
\frac{\partial}{\partial x_{\mu}} J_{\mu}=0
$$

### 4.7 Maxwell's equation using dual tensor

Using the electromagnetic tensor

$$
\left(F_{\mu \nu}\right)=\left(\begin{array}{cccc}
0 & B_{3} & -B_{2} & -\frac{i}{c} E_{1} \\
-B_{3} & 0 & B_{1} & -\frac{i}{c} E_{2} \\
B_{2} & -B_{1} & 0 & -\frac{i}{c} E_{3} \\
\frac{i}{c} E_{1} & \frac{i}{c} E_{2} & \frac{i}{c} E_{3} & 0
\end{array}\right)
$$

the dual tensor $G_{\mu \nu}$ is defined as

$$
G_{\mu \nu}=\frac{1}{2} \varepsilon_{\mu \nu \lambda \sigma} F_{\lambda \sigma}
$$

or

$$
\begin{aligned}
\left(G_{\mu \nu}\right) & =\left(\begin{array}{cccc}
0 & F_{34} & F_{42} & F_{23} \\
-F_{34} & 0 & F_{14} & F_{31} \\
-F_{42} & -F_{14} & 0 & F_{12} \\
-F_{23} & -F_{31} & -F_{12} & 0
\end{array}\right) \\
& =\left(\begin{array}{cccc}
0 & -\frac{i}{c} E_{3} & \frac{i}{c} E_{2} & B_{1} \\
\frac{i}{c} E_{3} & 0 & -\frac{i}{c} E_{1} & B_{2} \\
c & i \\
-\frac{i}{c} E_{2} & \frac{i}{c} E_{1} & 0 & B_{3} \\
-B_{1} & -B_{2} & -B_{3} & 0
\end{array}\right)
\end{aligned}
$$

Note that

$$
F_{\mu \nu}=\frac{1}{2} \varepsilon_{\mu \nu \lambda \sigma} G_{\lambda \sigma} .
$$

Using the Jacobi identity

$$
\frac{\partial F_{\mu v}}{\partial x_{\lambda}}+\frac{\partial F_{v \lambda}}{\partial x_{\mu}}+\frac{\partial F_{\lambda \mu}}{\partial x_{v}}=0
$$

we get

$$
\begin{aligned}
3 \frac{\partial G_{\mu v}}{\partial x_{v}} & =\varepsilon_{\mu \nu \lambda \sigma} \frac{\partial F_{\lambda \sigma}}{\partial x_{v}}+\varepsilon_{\mu \lambda \sigma v} \frac{\partial F_{\sigma v}}{\partial x_{\lambda}}+\varepsilon_{\mu \sigma v \lambda} \frac{\partial F_{v \lambda}}{\partial x_{\sigma}} \\
& =\varepsilon_{\mu \nu \lambda \sigma}\left(\frac{\partial F_{\lambda \sigma}}{\partial x_{v}}+\frac{\partial F_{\sigma v}}{\partial x_{\lambda}}+\frac{\partial F_{v \lambda}}{\partial x_{\sigma}}\right)=0
\end{aligned}
$$

since

$$
\varepsilon_{\mu \nu \lambda \sigma}=\varepsilon_{\mu \lambda \sigma v}=\varepsilon_{\mu \sigma \nu \lambda} .
$$

Then we have the Maxwell's equation,

$$
\frac{\partial G_{\mu \nu}}{\partial x_{v}}=0 .
$$

(a)

$$
\frac{\partial G_{1 v}}{\partial x_{v}}=\frac{\partial G_{11}}{\partial x_{1}}+\frac{\partial G_{12}}{\partial x_{2}}+\frac{\partial G_{13}}{\partial x_{3}}+\frac{\partial G_{14}}{\partial x_{4}}=-\frac{i}{c} \frac{\partial E_{3}}{\partial y}+\frac{i}{c} \frac{\partial E_{2}}{\partial z}+\frac{\partial B_{1}}{i c \partial t}=0
$$

or

$$
(\nabla \times \mathbf{E})_{1}=-\frac{\partial}{\partial t} B_{1}
$$

(b)

$$
\frac{\partial G_{2 v}}{\partial x_{v}}=\frac{\partial G_{21}}{\partial x_{1}}+\frac{\partial G_{22}}{\partial x_{2}}+\frac{\partial G_{23}}{\partial x_{3}}+\frac{\partial G_{24}}{\partial x_{4}}=\frac{i}{c} \frac{\partial E_{3}}{\partial x}-\frac{i}{c} \frac{\partial E_{1}}{\partial z}+\frac{\partial B_{2}}{i c \partial t}=0
$$

or

$$
(\nabla \times \mathbf{E})_{2}=-\frac{\partial}{\partial t} B_{2}
$$

(c)

$$
\frac{\partial G_{3 v}}{\partial x_{v}}=\frac{\partial G_{31}}{\partial x_{1}}+\frac{\partial G_{32}}{\partial x_{2}}+\frac{\partial G_{33}}{\partial x_{3}}+\frac{\partial G_{34}}{\partial x_{4}}=-\frac{i}{c} \frac{\partial E_{2}}{\partial x}+\frac{i}{c} \frac{\partial E_{1}}{\partial y}+\frac{\partial B_{3}}{i c \partial t}=0
$$

or

$$
(\nabla \times \mathbf{E})_{3}=-\frac{\partial}{\partial t} B_{3}
$$

(d)

$$
\frac{\partial G_{4 v}}{\partial x_{v}}=\frac{\partial G_{41}}{\partial x_{1}}+\frac{\partial G_{42}}{\partial x_{2}}+\frac{\partial G_{43}}{\partial x_{3}}+\frac{\partial G_{44}}{\partial x_{4}}=-\frac{\partial B_{1}}{\partial x}-\frac{\partial B_{2}}{\partial y}-\frac{\partial B_{3}}{\partial z}=0
$$

or

$$
\nabla \cdot \mathbf{B}=0
$$

Note

$$
G_{\mu \nu} F_{\mu \nu}=-\frac{4 i}{c}\left(B_{1} E_{1}+B_{2} E_{2}+B_{3} E_{3}\right)=-\frac{4 i}{c} \mathbf{E} \cdot \mathbf{B}
$$

### 4.8 Summary

The Maxwell's equation can be expressed by

$$
\frac{\partial F_{\mu v}}{\partial x_{v}}=\mu_{0} J_{\mu},
$$

and

$$
\frac{\partial G_{\mu v}}{\partial x_{v}}=0
$$

using the tensors $F$ and $G$.
5. Vector potential under the Lorentz transformation

$$
\begin{aligned}
& A_{\mu}=\left(\mathbf{A}, i \frac{1}{c} \phi\right) \\
& A_{\mu}^{\prime}=\left(\mathbf{A}^{\prime}, i \frac{1}{c} \phi^{\prime}\right) \\
& A_{\mu}^{\prime}=a_{\mu \nu} A_{\nu} \\
& A_{\mu}=\left(a^{-1}\right)_{\mu \nu} A_{\nu}^{\prime}
\end{aligned}
$$

$A_{1}{ }^{\prime}=\frac{c A_{1}-\beta \phi}{c \sqrt{1-\beta^{2}}}$
$A_{1}^{\prime}=\frac{c A_{1}-\beta \phi}{c \sqrt{1-\beta^{2}}}$
$A_{2}{ }^{\prime}=A_{2}$
$A_{2}{ }^{\prime}=A_{2}$
$A_{3}{ }^{\prime}=A_{3}$
$A_{3}{ }^{\prime}=A_{3}$
$A_{4}^{\prime}=-\frac{i}{c} \frac{\left(c \beta A_{1}-\phi\right)}{\sqrt{1-\beta^{2}}}$
$\phi^{\prime}=\frac{\phi-c \beta A_{1}}{\sqrt{1-\beta^{2}}}$
and
$A_{1}=\frac{c A_{1}^{\prime}+\beta \phi^{\prime}}{c \sqrt{1-\beta^{2}}}$
$A_{1}=\frac{c A_{1}{ }^{\prime}+\beta \phi^{\prime}}{c \sqrt{1-\beta^{2}}}$
$A_{2}=A_{2}{ }^{\prime}$
$A_{2}=A_{2}{ }^{\prime}$
$A_{3}=A_{3}{ }^{\prime}$
$A_{3}=A_{3}{ }^{\prime}$
$A_{4}=\frac{i}{c} \frac{\left(c \beta A_{1}^{\prime}+\phi^{\prime}\right)}{\sqrt{1-\beta^{2}}}$
$\phi=\frac{\left(c \beta A_{1}^{\prime}+\phi^{\prime}\right)}{\sqrt{1-\beta^{2}}}$

## 6. $\quad E$ and $B$ under the Lorentz transformation

### 6.1 Transformation

$$
\begin{aligned}
& F^{\prime}{ }_{\mu \nu}=a_{\mu \lambda} a_{v \sigma} F_{\lambda \sigma} \\
& a_{\mu \xi} a_{\nu \eta} F^{\prime}{ }_{\mu \nu}=a_{\mu \xi} a_{\nu \eta} a_{\mu \lambda} a_{v \sigma} F_{\lambda \sigma}=a_{\mu \xi} a_{\mu \lambda} a_{\nu \eta} a_{v \sigma} F_{\lambda \sigma}=\delta_{\xi \lambda} \delta_{\eta \sigma} F_{\lambda \sigma}=F_{\xi \eta}
\end{aligned}
$$

or

$$
F_{\mu \nu}=a_{\lambda \mu} a_{\sigma \nu} F_{\lambda \sigma}^{\prime}=\left(a^{T}\right)_{\mu \lambda}\left(a^{T}\right)_{\nu \sigma} F_{\lambda \sigma}^{\prime}=\left(a^{-1}\right)_{\mu \lambda}\left(a^{-1}\right)_{\nu \sigma} F_{\lambda \sigma}^{\prime}
$$

$E_{1}{ }^{\prime}=E_{1} \quad E_{1}{ }^{\prime}=E_{1}$
$E_{2}{ }^{\prime}=\gamma\left(E_{2}-c \beta B_{3}\right) \quad E_{2}{ }^{\prime}=\gamma(\mathbf{E}+\mathbf{v} \times \mathbf{B})_{2}$
$E_{3}{ }^{\prime}=\gamma\left(E_{3}+c \beta B_{2}\right) \quad E_{3}{ }^{\prime}=\gamma(\mathbf{E}+\mathbf{v} \times \mathbf{B})_{3}$
$B_{1}{ }^{\prime}=B_{1}$
$B_{1}{ }^{\prime}=B_{1}$
$B_{2}{ }^{\prime}=\gamma\left(B_{2}+\frac{\beta}{c} E_{3}\right)$
$B_{2}{ }^{\prime}=\gamma\left(\mathbf{B}-\frac{1}{c^{2}} \mathbf{v} \times \mathbf{E}\right)_{2}$
$B_{3}{ }^{\prime}=\gamma\left(B_{3}-\frac{\beta}{c} E_{2}\right)$
$B_{3}{ }^{\prime}=\gamma\left(\mathbf{B}-\frac{1}{c} \mathbf{v} \times \mathbf{E}\right)_{3}$

$$
\begin{array}{ll}
E_{1}=E_{1}{ }^{\prime} & E_{1}=E_{1}{ }^{\prime} \\
E_{2}=\gamma\left(E_{2}{ }^{\prime}+c \beta B_{3}{ }^{\prime}\right) & E_{2}=\gamma\left(\mathbf{E}^{\prime}-\mathbf{v} \times \mathbf{B}^{\prime}\right)_{2} \\
E_{3}=\gamma\left(E_{3}{ }^{\prime}-c \beta B_{2}{ }^{\prime}\right) & E_{3}=\gamma\left(\mathbf{E}^{\prime}-\mathbf{v} \times \mathbf{B}^{\prime}\right)_{3} \\
B_{1}=B_{1}{ }^{\prime} & B_{1}=B_{1}{ }^{\prime} \\
B_{2}=\gamma\left(B_{2}{ }^{\prime}-\frac{\beta}{c} E_{3}{ }^{\prime}\right) & B_{2}=\gamma\left(\mathbf{B}^{\prime}+\frac{1}{c^{2}} \mathbf{v} \times \mathbf{E}^{\prime}\right)_{2} \\
B_{3}=\gamma\left(B_{3}{ }^{\prime}+\frac{\beta}{c} E_{2}{ }^{\prime}\right) & B_{3}=\gamma\left(\mathbf{B}^{\prime}+\frac{1}{c^{2}} \mathbf{v} \times \mathbf{E}^{\prime}\right)_{3}
\end{array}
$$

### 6.2 Choice of the frame $S$ ' which has pure electric or pure magnetic fields

From the Sec.3.5, we find that
(1) $\mathbf{B}^{2}-\frac{1}{c^{2}} \mathbf{E}^{2}=$ invariant under the Lorentz transformation
(2) $\mathbf{E} \cdot \mathbf{B}=$ invariant under the Lorentz transformation

Here we assume that $\mathbf{E} \cdot \mathbf{B}=0$ and $\mathbf{B}^{2}-\frac{1}{c^{2}} \mathbf{E}^{2} \neq 0$
Then one can find a frame $S^{\prime}$ in which $\left(\boldsymbol{E}^{\prime}=0\right.$ and $\left.\boldsymbol{B}^{\prime} \neq 0\right)$ [pure magnetic field], or $\left(\boldsymbol{B}^{\prime}=\right.$ 0 and $\boldsymbol{E}^{\prime} \neq 0$ ) [pure electric field]. The proof is given in the following.
(a) Pure magnetic field $\left(E^{\prime}=0\right)$

We assume that $\boldsymbol{E}^{\prime}=0$. From the Lorentz transformation, we have

$$
\begin{array}{lll}
E_{1}^{\prime}=E_{1}=0 & E_{1}=0 \\
E_{2}=\gamma\left(E_{2}-c \beta B_{3}\right)=0 & \text { or } & E_{2}=c \beta B_{3}=v B_{3} \\
E_{3}^{\prime}=\gamma\left(E_{3}+c \beta B_{2}\right)=0 & & E_{3}=-c \beta B_{2}=-v B_{2}
\end{array}
$$

The condition $\mathbf{E} \cdot \mathbf{B}=0$ is satisfied since

$$
\mathbf{E} \cdot \mathbf{B}=E_{1} B_{1}+E_{2} B_{2}+E_{3} B_{3}=v B_{2} B_{3}-v B_{2} B_{3}=0
$$

The condition $\mathbf{B}^{2}-\frac{1}{c^{2}} \mathbf{E}^{2}=\mathbf{B}^{\prime 2}-\frac{1}{c^{2}} \mathbf{E}^{\prime 2} \neq 0$ can be rewritten as

$$
\mathbf{B}^{2}-\frac{1}{c^{2}} \mathbf{E}^{2}=\mathbf{B}^{\prime 2}>0
$$

This implies that one can find the frame where $\mathbf{B}^{\prime 2} \neq 0$ and $\boldsymbol{E}^{\prime}=0$.
((Note))
From the relation

$$
\begin{aligned}
& E_{1}=0 \\
& E_{2}=c \beta B_{3}=v B_{3} \\
& E_{3}=-c \beta B_{2}=-v B_{2}
\end{aligned}
$$

we get

## $\mathbf{E}=-\mathbf{v} \times \mathbf{B}$

(b) Pure electric field $\left(B^{\prime}=0\right)$

Next we assume that $\boldsymbol{B}^{\prime}=0$. Then we have

$$
\begin{array}{lll}
B_{1}{ }^{\prime}=B_{1}=0 & B_{1}=0 \\
B_{2}{ }^{\prime}=\gamma\left(B_{2}+\frac{\beta}{c} E_{3}\right)=0, & \text { or } & B_{2}=-\frac{v}{c^{2}} E_{3} \\
B_{3}{ }^{\prime}=\gamma\left(B_{3}-\frac{\beta}{c} E_{2}\right)=0 & & B_{3}=\frac{v}{c^{2}} E_{2}
\end{array}
$$

The condition $\mathbf{E} \cdot \mathbf{B}=0$ is satisfied since

$$
\mathbf{E} \cdot \mathbf{B}=E_{1} B_{1}+E_{2} B_{2}+E_{3} B_{3}=-\frac{v}{c^{2}} E_{2} E_{3}+\frac{v}{c^{2}} E_{2} E_{3}=0
$$

The condition $\mathbf{B}^{2}-\frac{1}{c^{2}} \mathbf{E}^{2}=\mathbf{B}^{\prime 2}-\frac{1}{c^{2}} \mathbf{E}^{\prime 2} \neq 0$ can be rewritten as

$$
\mathbf{B}^{2}-\frac{1}{c^{2}} \mathbf{E}^{2}=-\frac{1}{c^{2}} \mathbf{E}^{\prime 2}<0
$$

This implies that one can find the frame where $\mathbf{E}^{\prime 2} \neq 0$ and $\boldsymbol{B}^{\prime}=0$.
((Note))
From the relation

$$
\begin{aligned}
& B_{1}=0 \\
& B_{2}=-\frac{v}{c^{2}} E_{3} \\
& B_{3}=\frac{v}{c^{2}} E_{2}
\end{aligned}
$$

we get

$$
\mathbf{B}=\frac{1}{c^{2}}(\mathbf{v} \times \mathbf{E})
$$

## 7. Energy-momentum tensor and Maxwell's stress

## 7.1 force density

We define the vector of the force density as $f_{\mu}$

$$
F_{\mu \nu} J_{\nu}=f_{\mu}
$$

Here we have

$$
f_{i}=\rho E_{i}+(\mathbf{J} \times \mathbf{B})_{i}
$$

where

$$
\begin{aligned}
& \mathbf{J} \times \mathbf{B}=\left|\begin{array}{lll}
\hat{x} & \hat{y} & \hat{z} \\
J_{1} & J_{2} & J_{3} \\
B_{1} & B_{2} & B_{3}
\end{array}\right| \\
&\left\{\begin{array}{l}
f_{1} \\
\left\{\begin{array}{l}
E_{1}+\left(J_{2} B_{3}-J_{3} B_{2}\right) \\
f_{2}
\end{array}\right. \\
f_{3}
\end{array}=\rho E_{2}+\left(J_{3} B_{1}-J_{1} B_{3}\right)\right. \\
& f_{1}\left.=F_{1 v} J_{2}-J_{2} B_{1}\right) \\
&=B_{11} J_{1}+F_{12} J_{2}+B_{2} J_{3}-\frac{i}{c} E_{1}(i c \rho) \\
&=(\mathbf{B} \times \mathbf{J})_{1}+\rho E_{14} J_{4} \\
& f_{2}=(\mathbf{B} \times \mathbf{J})_{2}+\rho E_{2} \\
& f_{3}=(\mathbf{B} \times \mathbf{J})_{3}+\rho E_{3} \\
& f_{4}=F_{4 v} J_{v}=\frac{i}{c} E_{1} J_{1}+\frac{i}{c} E_{2} J_{2}+\frac{i}{c} E_{3} J_{3} \\
&=\frac{i}{c}(\mathbf{E} \cdot \mathbf{J})=i\left(\frac{\mathbf{E} \cdot \mathbf{J}}{c}\right)
\end{aligned}
$$

### 7.2 Maxwell's equation

The Maxwell's equation is given by

$$
\frac{\partial F_{v \lambda}}{\partial x_{\lambda}}=\mu_{0} J_{v}
$$

The current density:

$$
\begin{aligned}
& J_{\mu}=(\mathbf{J}, i c \rho) \\
& f_{\mu}=F_{\mu \nu} J_{v}=\frac{1}{\mu_{0}} F_{\mu \nu} \frac{\partial F_{v \lambda}}{\partial x_{\lambda}} \\
& \mu_{0} f_{\mu}=F_{\mu \nu} \frac{\partial F_{v \lambda}}{\partial x_{\lambda}}
\end{aligned}
$$

The left-hand side can be split into two terms,

$$
\mu_{0} f_{\mu}=\frac{\partial}{\partial x_{\lambda}}\left(F_{\mu \nu} F_{\nu \lambda}\right)-F_{\nu \lambda} \frac{\partial}{\partial x_{\lambda}} F_{\mu \nu}
$$

The second term:

$$
F_{\nu \lambda} \frac{\partial}{\partial x_{\lambda}} F_{\mu \nu}=\frac{1}{2} F_{\nu \lambda} \frac{\partial}{\partial x_{\lambda}} F_{\mu \nu}+\frac{1}{2} F_{\lambda \nu} \frac{\partial}{\partial x_{v}} F_{\mu \lambda}=\frac{1}{2} F_{\nu \lambda} \frac{\partial}{\partial x_{\lambda}} F_{\mu \nu}+\frac{1}{2} F_{\nu \lambda} \frac{\partial}{\partial x_{v}} F_{\lambda \mu}
$$

or

$$
F_{\nu \lambda} \frac{\partial}{\partial x_{\lambda}} F_{\mu \nu}=\frac{1}{2} F_{\nu \lambda}\left(\frac{\partial}{\partial x_{\lambda}} F_{\mu \nu}+\frac{\partial}{\partial x_{v}} F_{\lambda \mu}\right)=-\frac{1}{2} F_{\nu \lambda} \frac{\partial}{\partial x_{\mu}} F_{\nu \lambda}=-\frac{1}{4} \frac{\partial}{\partial x_{\mu}}\left(F_{\nu \lambda} F_{\nu \lambda}\right)
$$

Here we use the Jacobi identity;

$$
\frac{\partial}{\partial x_{\lambda}} F_{\mu \nu}+\frac{\partial}{\partial x_{\mu}} F_{\nu \lambda}+\frac{\partial}{\partial x_{\nu}} F_{\lambda \mu}=0 \quad \text { (Jacobi identity) }
$$

Then we have

$$
F_{\nu \lambda} \frac{\partial}{\partial x_{\lambda}} F_{\mu \nu}=-\frac{1}{4} \delta_{\mu \lambda} \frac{\partial}{\partial x_{\lambda}}\left(F_{\sigma \tau} F_{\sigma \tau}\right)
$$

The force density is rewritten as

$$
f_{\mu}=\frac{1}{\mu_{0}} \frac{\partial}{\partial x_{\lambda}}\left(F_{\mu \nu} F_{\nu \lambda}+\frac{1}{4} \delta_{\mu \lambda} F_{\sigma \tau} F_{\sigma \tau}\right)=\frac{\partial T_{\mu \lambda}}{\partial x_{\lambda}}
$$

with the symmetric energy-momentum tensor (Maxwell's stress tensor)

$$
\begin{aligned}
& T_{\mu \nu}=\frac{1}{\mu_{0}}\left(F_{\mu \lambda} F_{\lambda \nu}+\frac{1}{4} \delta_{\mu \lambda} F_{\sigma \tau} F_{\sigma \tau}\right) \\
& \operatorname{Tr}\left[T_{\mu \nu}\right]=T_{\mu \mu}=\frac{1}{\mu_{0}}\left(F_{\mu \lambda} F_{\lambda \mu}+\frac{1}{4} F_{\sigma \tau} F_{\sigma \tau}\right)=0
\end{aligned}
$$

### 7.3 Conservation law

$$
\begin{aligned}
& \frac{\partial u}{\partial t}+\nabla \cdot \mathbf{S}=-\mathbf{E} \cdot \mathbf{J} \\
& u=\frac{1}{2}\left(\varepsilon_{0} \mathbf{E}^{2}+\frac{1}{\mu_{0}} \mathbf{B}^{2}\right) \\
& \varepsilon_{0} \mu_{0} \frac{\partial \mathbf{S}}{\partial t}+\mathbf{f}=(\nabla \cdot \ddot{\mathbf{T}})
\end{aligned}
$$

where

$$
\begin{array}{ll}
\mathbf{S}=\frac{1}{\mu_{0}}(\mathbf{E} \times \mathbf{B}) & : \text { pointing vector } \\
\mathbf{G}=\varepsilon_{0} \mu_{0} \mathbf{S}=\frac{1}{c^{2}} \mathbf{S} \quad: \text { momentum of the field } \\
\mathbf{f}=\rho \mathbf{E}+(\mathbf{J} \times \mathbf{B}) & \\
T_{i j}=\left(\varepsilon_{0} E_{i} E_{j}+\frac{1}{\mu_{0}} B_{i} B_{j}\right)-\frac{1}{2} \delta_{i j}\left(\varepsilon_{0} \mathbf{E}^{2}+\frac{1}{\mu_{0}} \mathbf{B}^{2}\right)
\end{array}
$$

or

$$
\mu_{0} T_{i j}=\left(\frac{1}{c^{2}} E_{i} E_{j}+B_{i} B_{j}\right)-\frac{1}{2} \delta_{i j}\left(\frac{1}{c^{2}} \mathbf{E}^{2}+\mathbf{B}^{2}\right)
$$

where $c^{2}=\frac{1}{\varepsilon_{0} \mu_{0}}$

$$
\left(J_{\mu}\right)=(\mathbf{J}, i c \rho)
$$

$$
\begin{aligned}
& \left(F_{\mu \nu}\right)=\left(\begin{array}{cccc}
0 & B_{3} & -B_{2} & -\frac{i}{c} E_{1} \\
-B_{3} & 0 & B_{1} & -\frac{i}{c} E_{2} \\
B_{2} & -B_{1} & 0 & -\frac{i}{c} E_{3} \\
\frac{i}{c} E_{1} & \frac{i}{c} E_{2} & \frac{i}{c} E_{3} & 0
\end{array}\right) \\
& \mu_{0} T_{\mu \nu}=F_{\mu \alpha} F_{\alpha \nu}+\frac{1}{4} \delta_{\mu \nu} F_{\alpha \beta} F_{\alpha \beta}
\end{aligned}
$$

where

$$
\begin{aligned}
& F_{\alpha \beta} F_{\alpha \beta}=2\left[\left(B_{1}^{2}+B_{2}^{2}+B_{3}^{2}\right)-\frac{1}{c^{2}}\left(E_{1}^{2}+E_{2}^{2}+E_{3}^{2}\right)\right] \\
& F_{1 \alpha} F_{\alpha 2}=\frac{1}{c^{2}} E_{1} E_{2}+B_{1} B_{2} \\
& F_{2 \alpha} F_{\alpha 3}=\frac{1}{c^{2}} E_{2} E_{3}+B_{2} B_{3} \\
& F_{1 \alpha} F_{\alpha 3}=\frac{1}{c^{2}} E_{3} E_{1}+B_{3} B_{1} \\
& F_{1 \alpha} F_{\alpha 4}=-\frac{i}{c}\left(E_{2} B_{3}-B_{2} E_{3}\right)=-\frac{i \mu_{0}}{c} S_{1} \\
& F_{2 \alpha} F_{\alpha 4}=-\frac{i}{c}\left(E_{3} B_{1}-B_{1} E_{3}\right)=-\frac{i \mu_{0}}{c} S_{2} \\
& F_{3 \alpha} F_{\alpha 4}=-\frac{i}{c}\left(E_{1} B_{2}-B_{2} E_{1}\right)=-\frac{i \mu_{0}}{c} S_{3} \\
& F_{1 \alpha} F_{\alpha 1}=\frac{1}{c^{2}} E_{1}^{2}-\left(B_{2}^{2}+B_{3}^{2}\right) \\
& F_{2 \alpha} F_{\alpha 2}=\frac{1}{c^{2}} E_{2}^{2}-\left(B_{3}^{2}+B_{1}^{2}\right) \\
& F_{3 \alpha} F_{\alpha 3}=\frac{1}{c^{2}} E_{3}^{2}-\left(B_{1}^{2}+B_{2}^{2}\right) \\
& F_{4 \alpha} F_{\alpha 4}=\frac{1}{c^{2}}\left(E_{1}^{2}+E_{2}^{2}+E_{3}^{2}\right)
\end{aligned}
$$

The Maxwell's stress tensor is given by

$$
\begin{aligned}
& \mu_{0} T_{11}=\frac{1}{2 c^{2}}\left(E_{1}^{2}-E_{2}^{2}-E_{3}^{2}\right)+\frac{1}{2}\left(B_{1}^{2}-B_{2}^{2}-B_{3}^{2}\right) \\
& \mu_{0} T_{12}=\frac{1}{c^{2}} E_{1} E_{2}+B_{1} B_{2} \\
& \mu_{0} T_{13}=\frac{1}{c^{2}} E_{3} E_{1}+B_{3} B_{1} \\
& \mu_{0} T_{14}=-\frac{i}{c}\left(E_{2} B_{3}-B_{2} E_{3}\right)=-\frac{i \mu_{0}}{c} S_{1}=-i \mu_{0} c G_{1} \\
& \mu_{0} T_{22}=\frac{1}{2 c^{2}}\left(-E_{1}^{2}+E_{2}^{2}-E_{3}^{2}\right)+\frac{1}{2}\left(-B_{1}^{2}+B_{2}^{2}-B_{3}^{2}\right) \\
& \mu_{0} T_{23}=\frac{1}{c^{2}} E_{2} E_{3}+B_{2} B_{3} \\
& \mu_{0} T_{24}=-\frac{i}{c}\left(E_{3} B_{1}-B_{1} E_{3}\right)=-\frac{i \mu_{0}}{c} S_{2}=-i \mu_{0} c G_{2} \\
& \mu_{0} T_{33}=\frac{1}{2 c^{2}}\left(-E_{1}^{2}-E_{2}^{2}+E_{3}^{2}\right)+\frac{1}{2}\left(-B_{1}^{2}-B_{2}^{2}+B_{3}^{2}\right) \\
& \mu_{0} T_{34}=-\frac{i}{c}\left(E_{1} B_{2}-B_{2} E_{1}\right)=-\frac{i \mu_{0}}{c} S_{3}=-i \mu_{0} c G_{3} \\
& \mu_{0} T_{44}=\frac{1}{2}\left(B_{1}^{2}+B_{2}^{2}+B_{3}^{2}\right)+\frac{1}{2 c^{2}}\left(E_{1}^{2}+E_{2}^{2}+E_{3}^{2}\right)=\mu_{0} u
\end{aligned}
$$

Explicitly, the elements of $T$ are

$$
\begin{aligned}
& \left(T_{\mu \nu}\right)=\left(\begin{array}{cccc}
T_{11} & T_{12} & T_{13} & -i c G_{1} \\
T_{21} & T_{22} & T_{23} & -i c G_{2} \\
T_{31} & T_{33} & T_{34} & -i c G_{3} \\
-i c G_{1} & -i c G_{2} & -i c G_{3} & u
\end{array}\right) \\
& T_{11}+T_{22}+T_{33}=-u
\end{aligned}
$$

## 8. Lorentz force

### 8.1 Origin of the Lorentz force

Consider a particle of charge $q$ moving with velocity $\boldsymbol{v}$ (along the $x$ axis) with respect to the reference frame $S$ in a region with electric and magnetic fields $\boldsymbol{E}$ and $\boldsymbol{B}$.

In the frame $S$, the Lorentz force on this charge is given by

$$
\mathbf{F}=q(\mathbf{E}+\mathbf{v} \times \mathbf{B})=\left(q E_{1}, q\left(E_{2}-v B_{3}\right), q\left(E_{3}+v B_{2}\right)\right)
$$

In the frame $S^{\prime}$, the Lorentz force is given by

$$
\mathbf{F}^{\prime}=q \mathbf{E}^{\prime}=\left(q E_{1}^{\prime}, q E_{2}^{\prime}, q E_{3}^{\prime}\right)
$$

where $q$ is a relativistic invariant and is at rest.
The fields in $S$ and $S^{\prime}$ are related by

$$
\begin{aligned}
& E_{1}^{\prime}=E_{1} \\
& E_{2}^{\prime}=\gamma\left(E_{2}-v B_{3}\right) \\
& E_{3}^{\prime}=\gamma\left(E_{3}+v B_{2}\right)
\end{aligned}
$$

Then we have

$$
\begin{aligned}
& F_{1}^{\prime}=q E_{1}{ }^{\prime}=q E_{1} \\
& F_{2}{ }^{\prime}=q E_{2}^{\prime}=q \gamma\left(E_{2}-v B_{3}\right) \\
& F_{3}{ }^{\prime}=q E_{3}{ }^{\prime}=q \gamma\left(E_{3}+v B_{2}\right)
\end{aligned}
$$

What is the relation between $\boldsymbol{F}$ and $\boldsymbol{F}$ '?

$$
\begin{aligned}
& F_{1}^{\prime}=q E_{1}^{\prime}=q E_{1}=F_{1} \\
& F_{2}{ }^{\prime}=q E_{2}^{\prime}=q \gamma\left(E_{2}-v B_{3}\right)=\gamma F_{2} \\
& F_{3}^{\prime}=q E_{3}^{\prime}=q \gamma\left(E_{3}+v B_{2}\right)=\gamma F_{3}
\end{aligned}
$$

or

$$
\begin{aligned}
& F_{1}=F_{1}{ }^{\prime} \\
& \gamma F_{2}=F_{2}{ }^{\prime} \\
& \gamma F_{3}=F_{3}{ }^{\prime}
\end{aligned}
$$

## 8.2 force density and charge density

$$
\mathbf{f}=\rho \mathbf{E}+(\mathbf{J} \times \mathbf{B})
$$

We choose the frame $S^{\prime}$ in which the system with the charge density is at rest.
We now calculate the force density vector

$$
\mathbf{f}^{\prime}=\rho^{\prime} \mathbf{E}^{\prime}
$$

when $\mathbf{J}^{\prime}=0$ (the system is at rest).
We note the Lorentz transformation of 4-dimensional vector, current density and charge density

$$
\begin{aligned}
& J_{\mu}=(\mathbf{J}, i c \rho) \\
& J_{1}=\gamma\left(J_{1}{ }^{\prime}-i \beta J_{4}{ }^{\prime}\right)=\gamma\left(J_{1}{ }^{\prime}+\nu \rho^{\prime}\right) \\
& \rho=\gamma\left(\frac{\beta}{c} J_{1}{ }^{\prime}+\rho^{\prime}\right)
\end{aligned}
$$

Then we have

$$
\begin{aligned}
& \rho=\gamma \rho^{\prime} \\
& J_{1}=\gamma v \rho^{\prime}=\rho v
\end{aligned}
$$

The Lorentz transformation of $\boldsymbol{E}$ and $\boldsymbol{B}$,

$$
\begin{aligned}
& E_{1}^{\prime}=E_{1} \\
& E_{2}^{\prime}=\gamma\left(E_{2}-v B_{3}\right) \\
& E_{3}^{\prime}=\gamma\left(E_{3}+v B_{2}\right)
\end{aligned}
$$

Then we have

$$
\gamma \mathbf{f}^{\prime}=\left(\rho^{\prime} \gamma E_{1}, \rho^{\prime} \gamma^{2}\left(E_{2}-v B_{3}\right), \rho^{\prime} \gamma^{2}\left(E_{3}+v B_{2}\right)\right)
$$

or

$$
\gamma \mathbf{f}^{\prime}=\left(\rho E_{1}, \rho \gamma\left(E_{2}-v B_{3}\right), \rho \gamma\left(E_{3}+v B_{2}\right)\right)
$$

In the frame $S$, the Lorentz force is given by

$$
\mathbf{f}=\rho[\mathbf{E}+(\mathbf{v} \times \mathbf{B})]=\left(\rho E_{1}, \rho\left(E_{2}-v B_{3}\right), \rho\left(E_{3}+v B_{2}\right)\right.
$$

Thus we have

$$
\begin{aligned}
\gamma f_{1}^{\prime} & =f_{1} \\
f_{2} & =f_{2} \\
f_{3}^{\prime} & =f_{3}
\end{aligned}
$$

## 9. Lienard-Wiechert potential

### 9.1 Lienard-Wiechert potential

We now consider the Lienard-Wiechert potential


In the $S$ 'frame:

$$
\begin{array}{ll}
\phi^{\prime}=\frac{q}{4 \pi \varepsilon_{0}} \frac{1}{r^{\prime}} \\
\mathbf{A}^{\prime}=0 \\
A_{1}=\frac{A_{1}^{\prime}+\frac{v}{c^{2}} \phi^{\prime}}{\sqrt{1-\beta^{2}}} & A_{1}=\frac{\frac{v}{c^{2}} \phi^{\prime}}{\sqrt{1-\beta^{2}}} \\
A_{2}=A_{2}^{\prime} & A_{2}=0 \\
A_{3}=A_{3}^{\prime} & \text { or } \\
\phi=\frac{v A_{1}^{\prime}+\phi^{\prime}}{\sqrt{1-\beta^{2}}} & A_{3}=0 \\
& \phi=\frac{\phi^{\prime}}{\sqrt{1-\beta^{2}}}=\frac{q}{4 \pi \varepsilon_{0}} \frac{1}{\sqrt{1-\beta^{2}}} \frac{1}{r^{\prime}}
\end{array}
$$

Then we get

$$
\phi=\frac{\phi^{\prime}}{\sqrt{1-\beta^{2}}}=\frac{q}{4 \pi \varepsilon_{0}} \frac{1}{\sqrt{1-\beta^{2}}} \frac{1}{\sqrt{x_{1}^{\prime 2}+x_{2}{ }^{\prime 2}+x_{3}{ }^{\prime 2}}}
$$

where

$$
\begin{aligned}
& x_{1}{ }^{\prime}=\gamma\left(x_{1}-v t\right) \\
& x_{2}{ }^{\prime}=x_{2} \\
& x_{3}^{\prime}=x_{3}
\end{aligned}
$$

The scalar potential $\phi$ is given by
or

$$
\phi=\frac{q}{4 \pi \varepsilon_{0}} \frac{1}{R^{*}}
$$

with

$$
R^{*}=\sqrt{\left(x_{1}-v t\right)^{2}+\left(1-\beta^{2}\right)\left(x_{2}^{2}+x_{3}^{2}\right)}
$$

Similarly we have for the vector potential

$$
\mathbf{A}=\left(A_{1}, 0,0\right)
$$

with

$$
A_{1}=\frac{\frac{v}{c^{2}} \phi^{\prime}}{\sqrt{1-\beta^{2}}}=\frac{v}{c^{2}} \frac{q}{4 \pi \varepsilon_{0}} \frac{1}{R^{*}}=\frac{q v \mu_{0}}{4 \pi} \frac{1}{R^{*}}
$$

The electric field $\boldsymbol{E}$ and the magnetic field $\boldsymbol{B}$ are given by

$$
\begin{aligned}
& \mathbf{E}=-\frac{\partial}{\partial t} \mathbf{A}-\nabla \phi \quad \text { and } \quad \mathbf{B}=\nabla \times \mathbf{A} \\
& \mathbf{E}=\frac{q}{4 \pi \varepsilon_{0}}\left(1-\beta^{2}\right) \frac{\mathbf{R}}{R^{* 3}} \\
& \mathbf{B}=\frac{\mathbf{v}}{c^{2}} \times \mathbf{E}
\end{aligned}
$$

where

$$
\mathbf{R}=(x-v t, y, z)
$$

For a slow moving charge $(v \ll c)$, we can take for $\boldsymbol{E}$ the Coulomb field. Then whave

$$
\mathbf{B}=\frac{\mathbf{v}}{c^{2}} \times \mathbf{E}=\frac{q \mathbf{v} \times \mathbf{r}}{4 \pi \varepsilon_{0} c^{2} r^{2}}=\frac{\mu_{0}}{4 \pi} \frac{q \mathbf{v} \times \mathbf{r}}{r^{2}}
$$

## ((Mathematica-10))

Lienard-Wiechert potential

$$
\begin{aligned}
& \text { <<Calculus`VectorAnalysis` } \\
& \text { SetCoordinates [Cartesian [x,y,z]] } \\
& \text { Cartesian }[\mathrm{x}, \mathrm{y}, \mathrm{z}] \\
& \mathbf{R}=\sqrt{(\mathbf{X}-\mathbf{v} \mathbf{t})^{2}+\left(\mathbf{1}-\beta^{2}\right)\left(\mathbf{y}^{2}+\mathbf{z}^{\mathbf{2}}\right)} \\
& \sqrt{(-t v+x)^{2}+\left(y^{2}+z^{2}\right)\left(1-\beta^{2}\right)} \\
& \phi=\frac{\mathrm{q}}{4 \pi \epsilon 0} \frac{1}{\mathrm{R}} \\
& \frac{\mathrm{q}}{4 \pi \sqrt{\left.(-\mathrm{tv}+\bar{x})^{2--}+\mathrm{y}^{2}+\mathrm{z}^{2}\right)}\left(\overline{1}-\bar{\beta}^{2}\right) \in 0} \\
& A 1=\frac{q v \mu 0}{4 \pi} \frac{1}{R} \\
& \frac{q \mathrm{~V} \mu 0}{4 \pi \sqrt{\left.(-\mathrm{tV}+\mathrm{X})^{2--}+\mathrm{y}^{2--}+\mathrm{z}^{2}\right)\left(\overline{1}^{--\beta^{2}}\right)}} \\
& A=\{A 1,0,0\} \\
& \left\{\frac{q v \mu 0}{4 \pi \sqrt{(-t v+X)^{2--}\left(\bar{y}^{2}+\bar{z}^{2}\right)\left(1-\beta^{2}\right)}}, 0,0\right\} \\
& \text { B1=Curl [A] / /FullSimplify } \\
& \left\{0, \frac{q \vee z\left(-1+\beta^{2}\right) \mu 0}{4 \pi\left((-t v+X)^{2}-\left(\mathbf{y}^{2}+z^{2}\right)\left(-1+\beta^{2}\right)\right)^{3 / 2}}\right. \\
& \left.-\frac{q \vee y\left(-1+\beta^{2}\right) \mu 0}{4 \pi\left((-t v+X)^{2}-\left(y^{2}+z^{2}\right)\left(-1+\beta^{2}\right)\right)^{3 / 2}}\right\}
\end{aligned}
$$

Electric field in the frame $S$

$$
\mathrm{E} 1=-\operatorname{Grad}[\phi]-\mathrm{D}[\mathrm{~A}, \mathrm{t}] / .\left\{\mu 0 \rightarrow \mathbf{1} /\left(\epsilon 0 \mathrm{c}^{2}\right)\right\} / / \text { FullSimplify }
$$

$$
\begin{aligned}
& \left\{\frac{q\left(-c^{2}+v^{2}\right)(t v-x)}{4 c^{2} \pi\left((-t v+x)^{2}-\left(y^{2}+z^{2}\right)\left(-1+\beta^{2}\right)\right)^{3 / 2} \in 0},\right. \\
& -\frac{q y\left(-1+\beta^{2}\right)}{4 \pi\left((-t v+x)^{2}-\left(y^{2}+z^{2}\right)\left(-1+\beta^{2}\right)\right)^{3 / 2} \in 0}, \\
& \left.-\frac{q z\left(-1+\beta^{2}\right)}{4 \pi\left((-t v+x)^{2}-\left(y^{2}+z^{2}\right)\left(-1+\beta^{2}\right)\right)^{3 / 2} \in 0}\right\} \\
& \mathrm{V} 1=\{\mathrm{v}, 0,0\} \\
& \{\mathrm{v}, 0,0\} \\
& \text { eq1 }=\frac{1}{\mathbf{c}^{2}} \operatorname{Cross}[\text { V1, E1] // Simplify } \\
& \left\{0, \frac{q v z\left(-1+\beta^{2}\right)}{4 \mathrm{c}^{2} \pi\left((-\mathrm{tv}+\mathrm{x})^{2}-\left(\mathrm{y}^{2}+\mathrm{z}^{2}\right)\left(-1+\beta^{2}\right)\right)^{3 / 2} \in 0},\right. \\
& \left.-\frac{q v y\left(-1+\beta^{2}\right)}{4 \mathrm{c}^{2} \pi\left((-\mathrm{tv}+\mathrm{x})^{2}-\left(\mathrm{y}^{2}+\mathrm{z}^{2}\right)\left(-1+\beta^{2}\right)\right)^{3 / 2} \in 0}\right\} \\
& \text { eq1-B1 / }\left\{\mu 0 \rightarrow \frac{1}{\mathbf{c}^{2} \in 0}\right\} / / \text { Simplify } \\
& \{0,0,0\}
\end{aligned}
$$

### 9.2 Distribution of the electric field

$$
\mathbf{E}=\frac{q}{4 \pi \varepsilon_{0}}\left(1-\beta^{2}\right) \frac{\mathbf{R}_{p}}{R^{* 3}}
$$

where

$$
\begin{aligned}
& R^{*}=\sqrt{(x-v t)^{2}+\left(1-\beta^{2}\right)\left(y^{2}+z^{2}\right)} \\
& \mathbf{R}^{*}=\left(x-v t, \sqrt{1-\beta^{2}} y, \sqrt{1-\beta^{2}} z\right) \\
& \mathbf{R}_{p}=(x-v t, y, z)
\end{aligned}
$$

$\boldsymbol{R}_{\mathrm{p}}$ is the relative coordinate of the field point and the charge point. The electric field is along the position vector $\boldsymbol{R}_{\mathrm{p}}$. $\boldsymbol{R}_{\mathrm{p}}$ is a vector from the instantaneous location of the charge in $S$ to the point where $\boldsymbol{E}$ is measured in $S$.
((Mathematica-11))
The electric field of a charge moving with the constant speed with $v(\beta=v / c)$ on the unit circle of the real space

```
Lienard-Wiechert problem; field for a uniformly moving charge
<<Graphics PlotField`
    \(\left.E 1 X_{[\theta-,}, \beta_{-}\right]:=\frac{X\left(1-\beta^{2}\right)}{\left(\sqrt{\left.\bar{x}^{2}+\left(1-\beta^{2}\right) \bar{y}^{2}\right)^{3}}\right.} / .\{X \rightarrow \operatorname{Cos}[\theta], y \rightarrow \operatorname{Sin}[\theta]\} / /\) Simplify
    \(\left.E 1 Y_{[\theta-}, \beta\right]:=\frac{y\left(1-\beta^{2}\right)}{\left(\sqrt{x^{2}+\left(1-\beta^{2}\right) y^{2}}\right)^{3}} / .\{X \rightarrow \operatorname{Cos}[\theta], y \rightarrow \operatorname{Sin}[\theta]\} / /\) Simplify
    \(\operatorname{E1X}[\theta, \beta]\)
    \(\frac{\left(1-\beta^{2}\right) \operatorname{Cos}[\theta]}{\left(\operatorname{Cos}[\theta]^{2}-\left(-1+\beta^{2}\right) \operatorname{Sin}[\theta]^{2}\right)^{3 / 2}}\)
    \(\operatorname{E1Y}[\theta, \beta]\)
    \(\frac{\left(1-\beta^{2}\right) \operatorname{Sin}[\theta]}{\left(\operatorname{Cos}[\theta]^{2}-\left(-1+\beta^{2}\right) \operatorname{Sin}[\theta]^{2}\right)^{3 / 2}}\)
\(\operatorname{si}[\beta]\) : \(=\) Table \([\{\{E 1 X[\theta, \beta], E 1 Y[\theta, \beta]\},\{E 1 X[\theta, \beta], E 1 Y[\theta, \beta]\}\)
\(\},\{\theta, 0,2 \pi, \pi / 32\}]\)
s2[ \(\beta\) _]:=ListPlotVectorField[ Evaluate[s1[ \(\beta\) ]]//N//Chop,C
olorFunction \(\rightarrow\) Hue,
AspectRatio \(\rightarrow\) Automatic, ScaleFactor \(\rightarrow 1\), Frame \(\rightarrow\) True, PlotPoi
nts \(\rightarrow 20\), AxesOrigin \(\rightarrow\{0,0\}\), DeFaultColor \(\rightarrow\) Hue [0.6], DisplayF
unction \(\rightarrow\) Identity]
    General :: spell1 : Possible spelling error: new symbol
            name "DeFaultColor " is similar to existing symbol "DefaultColor ". More...
\(\beta=0,0.04,0.08,0.12,0.16\)
\(\beta=0.20,0.24,0.28,0.32,0.36\)
```

ps1=Evaluate [Table[s2[ $\beta$ ], $\{\beta, 0,0.36,0.04\}]$ ] Show[Graphic sArray[Partition[ps1,5]], DisplayFunction $\rightarrow \$$ DisplayFunct ion]










-GraphicsArray-
$\beta=0.4,0.42,0.44,0.46,0.48$
$\beta=0.50,0.52,0.54,0.56,0.58$
ps2=Evaluate[Table[s2[ $\beta$ ], $\{\beta, 0.4,0.58,0.02\}$ ]] Show [Graph icsArray[Partition[ps2,5]],DisplayFunction $\rightarrow$ \$DisplayFun ction]

ps3=Evaluate [Table [s2[ $\beta$ ], $\{\beta, 0.6,0.78,0.02\}$ ]] ; Show [Graph icsArray[Partition[ps3,5]],DisplayFunction $\rightarrow$ \$DisplayFun ction]

ps4=Evaluate[Table[s2[ $\beta$ ], $\{\beta, 0.8,0.89,0.01\}]$ ] Show[Graph icsArray[Partition[ps4,5]],DisplayFunction $\rightarrow$ \$DisplayFun ction]

ps5=Evaluate [Table[s2[ $\beta$ ], $\{\beta, 0.9,0.99,0.01\}]$ ] Show [Graph icsArray[Partition[ps5,5]],DisplayFunction $\rightarrow$ \$DisplayFun ction]

10. Relativity of Electric field and magnetic field

$E=0$ and $B \neq 0$
We consider the charge $q$ moving along the $x$ axis in the presence of the magnetic field $\boldsymbol{B}$ (the frame $S$ ). In the frame $S$, there is only an external magnetic field $\boldsymbol{B}$. Thus the magnetic force on the charge is given by

$$
\mathbf{F}_{\perp}=q(\mathbf{v} \times \mathbf{B})
$$

Suppose that there is no electric field $(\boldsymbol{E}=0)$ in the frame $S(\mathbf{B} \neq 0)$. The $\boldsymbol{E}^{\prime}$ and $\boldsymbol{B}^{\prime}$ in the frame $S^{\prime}$ are related to those in the frame $S$ as

$$
\begin{aligned}
& E_{1}^{\prime}=E_{1}=0 \\
& E_{2}{ }^{\prime}=\gamma\left(E_{2}-c \beta B_{3}\right)=-\gamma B_{3} \\
& E_{3}^{\prime}=\gamma\left(E_{3}+c \beta B_{2}\right)=\gamma B_{2} \\
& B_{1}^{\prime}=B_{1} \\
& B_{2}{ }^{\prime}=\gamma\left(B_{2}+\frac{\beta}{c} E_{3}\right)=\gamma B_{2} \\
& B_{3}{ }^{\prime}=\gamma\left(B_{3}-\frac{\beta}{c} E_{2}\right)=\gamma B_{3}
\end{aligned}
$$

or

$$
\begin{equation*}
\mathbf{E}^{\prime}=\gamma(\mathbf{v} \times \mathbf{B})=\mathbf{v} \times \mathbf{B}^{\prime} \tag{1}
\end{equation*}
$$

Then the force (electric force) on the charge $q$ in the frame $S^{\prime}$ is

$$
\mathbf{F}_{\perp}^{\prime}=q \mathbf{E}^{\prime}=q \gamma(\mathbf{v} \times \mathbf{B})
$$

since the charge $q$ is the same for any frame and the particle is at rest in the frame $\mathrm{S}^{\prime}$. There is no force due to $\boldsymbol{B}^{\prime}$ since the particle is at rest in the frame $S^{\prime} . \mathbf{F}_{\perp}{ }_{\perp}$ is the force of $\boldsymbol{F}^{\prime}$ ' in a direction perpendicular to the velocity $\boldsymbol{v}$. Thus we have

$$
\mathbf{F}_{\perp}=\frac{1}{\gamma} \mathbf{F}_{\perp}{ }^{\prime}
$$

## 11. Derivation of the Biot Savart law

$\boldsymbol{B}^{\prime}=0$ and $\boldsymbol{E}^{\prime} \neq 0$.
We consider that the magnetic field $\boldsymbol{B}^{\prime}=0$ in the frame $S^{\prime}$. In the frame $S^{\prime}$, there is only an external electric field $\boldsymbol{E}^{\prime}$ (the point charge is at rest). The $\boldsymbol{E}$ and $\boldsymbol{B}$ in the frame $S$ are related to those in the frame $S^{\prime}$ as

$$
\begin{array}{ll}
E_{1}=E_{1}^{\prime} & B_{1}=0 \\
E_{2}=\gamma E_{2}^{\prime} & B_{2}=-\frac{\gamma}{c^{2}} v E_{3}^{\prime} \\
E_{3}=\gamma E_{3}^{\prime} & B_{3}=\frac{\gamma}{c} \beta E_{2}^{\prime}=\frac{\gamma}{c^{2}} v E_{2}^{\prime}
\end{array}
$$

or

$$
\begin{equation*}
\mathbf{B}=\frac{\gamma}{c^{2}}\left(\mathbf{v} \times \mathbf{E}^{\prime}\right)=\frac{1}{c^{2}}(\mathbf{v} \times \mathbf{E}), \tag{2}
\end{equation*}
$$

Using the result from the Lienard-Wiechert potential $(\beta \ll 1)$ (see Sec.8)

$$
\begin{aligned}
& \mathbf{E}=\frac{q}{4 \pi \varepsilon_{0}}\left(1-\beta^{2}\right) \frac{\mathbf{R}}{R^{* 3}} \approx \frac{q}{4 \pi \varepsilon_{0}} \frac{\mathbf{r}}{r^{3}} \\
& \mathbf{B}=\frac{1}{c^{2}}(\mathbf{v} \times \mathbf{E})=\frac{1}{c^{2}} \frac{q}{4 \pi \varepsilon_{0}} \frac{\mathbf{v} \times \mathbf{r}}{r^{3}}=\frac{\mu_{0}}{4 \pi} \frac{q \mathbf{v} \times \mathbf{r}}{r^{3}}
\end{aligned}
$$

which is the application of the Biot-Savart law to a point charge.

## 12. Ampere's law (Feynman 13-9)

We consider that the electrons located on the linear chain (the line density $-\lambda_{0}$ ) moves at the velocity $\boldsymbol{v}$. At the same time there are positive ions located on the same chain (the line density $\lambda_{0}$ ). We now consider the frame $S^{\prime}$ which moves at the velocity $\boldsymbol{v}$.

((Formula))

$$
\rho=\gamma \rho^{\prime}
$$

where $\rho$ for the frame where the particle moves at the velocity $v$ along the $x$ axis, and $\rho^{\prime}$ for the frame where the particle is at rest.

We assume that
(1) The line densities of electrons and positive ions are given by $-\lambda_{0}$ and $+\lambda_{0}$ in the frame S .
(2) The line densities of electrons and positive ions are given by $\lambda_{-}$and $\lambda_{+}$in the frame $S$,

$$
\begin{array}{lll}
\left(-\lambda_{0}\right)=\gamma\left(-\lambda_{-}\right) & \text {or } & \lambda_{-}=\frac{1}{\gamma} \lambda_{0}=\sqrt{1-\frac{v^{2}}{c^{2}}} \lambda_{0} \quad \text { for electrons } \\
\lambda_{+}=\gamma \lambda_{0} & \text { or } & \lambda_{+}=\gamma \lambda_{0}=\frac{1}{\sqrt{1-\frac{v^{2}}{c^{2}}}} \lambda_{0} \quad \text { for ions }
\end{array}
$$

The net line charge density in the frame $S^{\prime}$ is

$$
\lambda^{\prime}=\lambda_{+}-\lambda_{-}=\frac{1}{\sqrt{1-\frac{v^{2}}{c^{2}}}} \lambda_{0}-\sqrt{1-\frac{v^{2}}{c^{2}}} \lambda_{0}=\frac{1}{\sqrt{1-\frac{v^{2}}{c^{2}}}} \frac{v^{2}}{c^{2}} \lambda_{0}=\gamma \frac{v^{2}}{c^{2}} \lambda_{0}
$$

((Note))
This relation can be also derived from the Lorentz transformation of the 4dimensional current density

$$
\begin{aligned}
& J_{\mu}=(\mathbf{J}, i c \rho)=(\rho \mathbf{v}, i c \rho) \\
& J_{\mu}{ }^{\prime}=a_{\mu \nu} J_{v} \\
& J_{1}{ }^{\prime}=\gamma\left(J_{1}+i \beta J_{4}\right)=\gamma\left(J_{1}-v \rho\right) \\
& \rho^{\prime}=\gamma\left(-\frac{\beta}{c} J_{1}+\rho\right) \\
& J_{\mu}=a_{v \mu} J_{v}{ }^{\prime} \\
& J_{1}=\gamma\left(J_{1}{ }^{\prime}-i \beta J_{4}{ }^{\prime}\right)=\gamma\left(J_{1}{ }^{\prime}+v \rho^{\prime}\right) \\
& \rho=\gamma\left(\frac{\beta}{c} J_{1}{ }^{\prime}+\rho^{\prime}\right)
\end{aligned}
$$

Here we define

$$
\begin{aligned}
& \lambda=A \rho \\
& \lambda^{\prime}=A \rho^{\prime}
\end{aligned}
$$

A is the same for the $S$ and $S^{\prime}$, since the plane of $A$ is perpendicular to $\boldsymbol{v}$.

$$
\lambda^{\prime}=\gamma\left(-\frac{\beta}{c} J_{1}+\lambda\right)=\gamma \frac{\beta}{c} \lambda_{0} v=\gamma \frac{v^{2}}{c^{2}} \lambda_{0}
$$

where $I_{1}=A J_{1}=\left(-\lambda_{0}\right) v$ and $\lambda=0$.
So the positive line density produces an electric field $\boldsymbol{E}$ '. We use the Gauss's law.


The electric field $\boldsymbol{E}$ ' at the distance $s$ from the axis of the cylinder,

$$
E^{\prime}(2 \pi \mathrm{~s} h)=\frac{1}{\varepsilon_{0}}\left(h \lambda^{\prime}\right)
$$

where $s$ is the radius of the Gaussian surface (cylinder).
or

$$
E^{\prime}=\frac{\lambda^{\prime}}{2 \pi \varepsilon_{0} s}=\frac{1}{2 \pi \varepsilon_{0} s} \gamma \frac{v^{2}}{c^{2}} \lambda_{0}
$$

So there is an electrical force on $q$ in $S^{\prime}$;

$$
F_{\perp}^{\prime}=q E^{\prime}=\frac{q}{2 \pi \varepsilon_{0} s} \gamma \frac{v^{2}}{c^{2}} \lambda_{0} .
$$

But if there is a force on the test charge $q$ in $S^{\prime}$, there must be one in $S$. In fact, one can calculate it by using the transformation rules for forces. Since $q$ is at rest $S^{\prime}$ and $F_{\perp}$ is perpendicular to the $x$ axis. Then we have

$$
F_{\perp}{ }^{\prime}=\gamma F_{\perp} \quad \text { or } \quad F_{\perp}=\frac{1}{\gamma} F_{\perp}{ }^{\prime}
$$

Using this result we have

$$
F_{\perp}=\frac{1}{\gamma} F_{\perp}{ }^{\prime}=\frac{q}{2 \pi \varepsilon_{0} s} \frac{v^{2}}{c^{2}} \lambda_{0}=\frac{q \mu_{0}}{2 \pi \mathrm{~s}} v\left(v \lambda_{0}\right)=q \frac{\mu_{0}\left(v \lambda_{0}\right)}{2 \pi \mathrm{~s}} v
$$

where $B=\frac{\mu_{0}\left(\nu \lambda_{0}\right)}{2 \pi \mathrm{~s}}$ is a magnetic field due to the line current density $v \lambda_{0}$ (Ampere's law). The force has a form as $F=q v B$.

## 13. Derivation of the Ampere's law from relativity

We analyze the fields and currents as viewed from two frames. $S$ where the ions are at rest. $S^{\prime}$ where the electrons are, on the average, at rest.

$$
\begin{aligned}
& J_{\mu}=(\mathbf{J}, i c \rho) \\
& J_{\mu}=\left(a^{-1}\right)_{\mu \nu} J_{v}{ }^{\prime}=a_{v \mu} J_{v}{ }^{\prime}
\end{aligned}
$$

Multiplying the cross-sectional area $(A)$ of the wires, we obtain the following transformation for currents and linear charge densities.

$$
\begin{aligned}
& I_{\mu}=A J_{\mu}=(A \mathbf{J}, i c A \rho)=(\mathbf{I}, i c \lambda) \\
& I_{\mu}=a_{v \mu} I_{\nu}{ }^{\prime} \\
& I_{1}=\gamma\left(I_{1}{ }^{\prime}-i \beta I_{4}{ }^{\prime}\right)=\gamma\left(I_{1}{ }^{\prime}+v \lambda^{\prime}\right) \\
& I_{4}=i c \lambda=\gamma\left(i \beta I_{1}{ }^{\prime}+I_{4}{ }^{\prime}\right)=\gamma\left(i \frac{v}{c} I_{1}{ }^{\prime}+i c \lambda^{\prime}\right)
\end{aligned}
$$

or

$$
\begin{aligned}
& I_{ \pm}=\gamma\left(I_{ \pm}^{\prime}+v \lambda_{ \pm}^{\prime}\right) \\
& \lambda_{ \pm}=\gamma\left(\frac{v}{c^{2}} I_{ \pm}^{\prime}+\lambda_{ \pm}^{\prime}\right)
\end{aligned}
$$

where $\lambda=A \rho$, the subscript 1 is neglected and the plus and minus subscript refer to the ions and the electrons, respectively.

In $S^{\prime}$ we know that $I_{-}^{\prime}=0$ since the electrons are at rest.

$$
\lambda_{-}=\gamma\left(\frac{v}{c^{2}} I_{-}{ }^{\prime}+\lambda_{-}{ }^{\prime}\right)=\gamma \lambda_{-}{ }^{\prime}
$$

In $S$ the net charge per unit length must vanish.

$$
0=\lambda_{+}+\lambda_{-}=\lambda_{+}+\gamma \lambda_{-}^{\prime}
$$

or

$$
\lambda_{-}^{\prime}=-\frac{\lambda_{+}}{\gamma}
$$

The fields in $S^{\prime}$ due to $\lambda_{-}{ }^{\prime}$ are

$$
\begin{aligned}
& \mathbf{E}_{-}^{\prime}=\frac{\lambda_{-}^{\prime}}{2 \pi \varepsilon_{0} r^{\prime}} \mathbf{e}_{r} \\
& \mathbf{B}_{-}^{\prime}=0
\end{aligned}
$$



The fields in $S$ due to $\lambda_{+}$are

$$
\begin{aligned}
& \mathbf{E}_{+}=\frac{\lambda_{+}}{2 \pi \varepsilon_{0} r} \mathbf{e}_{r} \\
& \mathbf{B}_{+}=0
\end{aligned}
$$

We now consider the field transformation for

$$
\mathbf{E}_{-}^{\prime}=\frac{\lambda_{-}^{\prime}}{2 \pi \varepsilon_{0} r^{\prime}} \mathbf{e}_{r}
$$

$$
\mathbf{B}_{-}{ }^{\prime}=0
$$

noting that $\hat{r}=\hat{r}^{\prime}$ (perpendicular to the $x$ axis), we find that the fields in $S$ are

$$
\begin{array}{ll}
E_{1}=E_{1}{ }^{\prime}=0 & B_{1}=B_{1}{ }^{\prime}=0 \\
E_{2}=\gamma\left(E_{2}{ }^{\prime}+c \beta B_{3}{ }^{\prime}\right)=\gamma E_{2}{ }^{\prime} & B_{2}=\frac{\gamma}{c}\left(c B_{2}{ }^{\prime}-\beta E_{3}{ }^{\prime}\right)=-\frac{\gamma}{c^{2}} E_{3}{ }^{\prime} \\
E_{3}=\gamma E_{3}{ }^{\prime} & B_{3}=\frac{\gamma}{c}\left(\beta E_{2}{ }^{\prime}+c B_{3}{ }^{\prime}\right)=\frac{\gamma}{c^{2}} E_{2}{ }^{\prime}
\end{array}
$$

$$
\begin{aligned}
& \mathbf{E}_{-}=\gamma \frac{\mathbf{e}_{r}}{2 \pi \varepsilon_{0} r} \lambda_{-}{ }^{\prime}, \\
& \mathbf{B}_{-}=\frac{\gamma}{c^{2}}\left(\mathbf{v} \times \mathbf{E}_{-}{ }^{\prime}\right)
\end{aligned}
$$

Then the total fields in the frame $S$ are

$$
\begin{aligned}
& \mathbf{E}=\mathbf{E}_{+}+\mathbf{E}_{-}=\frac{\mathbf{e}_{r}}{2 \pi \varepsilon_{0} r}\left(\lambda_{+}+\gamma \lambda_{-}^{\prime}\right)=0 \\
& \mathbf{B}=\mathbf{B}_{+}+\mathbf{B}_{-}=\mathbf{B}_{-}=\frac{\gamma}{c^{2}}\left(\mathbf{v} \times \mathbf{E}_{-}^{\prime}\right)=\frac{\gamma}{c^{2}}\left(\mathbf{v} \times \mathbf{e}_{r}\right) \frac{\lambda_{-}^{\prime}}{2 \pi \varepsilon_{0} r^{\prime}}=\frac{\gamma \lambda_{-}{ }^{\prime} v}{2 \pi \varepsilon_{0} c^{2}} \frac{\left(\mathbf{e}_{x} \times \mathbf{e}_{r}\right)}{r}
\end{aligned}
$$

Since $I_{-}=\gamma\left(I_{-}{ }^{\prime}+\nu \lambda_{-}{ }^{\prime}\right)$ and $I_{-}{ }^{\prime}=0$, we have

$$
I_{-}=\gamma \nu \lambda_{-}^{\prime}
$$

Using $\mathbf{e}_{x} \times \mathbf{e}_{r}=\mathbf{e}_{\phi}$, we obtain

$$
\begin{aligned}
& \mathbf{B}=\frac{\mu_{0} I_{-}}{2 \pi r} \mathbf{e}_{\phi} \\
& \mathbf{E}=0
\end{aligned}
$$

We see that a magnetic field due to current flow is a relativistic effect.
14. Capacitance moving along the $x$ axis with a uniform velocity
14.1 The capacitance moves along the $x$ direction which is parallel to the electric field of the capacitance.


In the frame $S$ ' where the charges are at rest.

$$
\begin{aligned}
& E_{1}{ }^{\prime}=\frac{\sigma^{\prime}}{\varepsilon_{0}} \\
& E_{2}=0 \\
& E_{3}{ }^{\prime}=0
\end{aligned}
$$

and

$$
\begin{aligned}
& B_{1}{ }^{\prime}=0 \\
& B_{2}=0 \\
& B_{3}=0
\end{aligned}
$$

$$
\begin{array}{ll}
E_{1}=E_{1}{ }^{\prime} & B_{1}=0 \\
E_{2}=0 & B_{2}=0 \\
E_{3}=0 & B_{3}=0
\end{array}
$$

Thus we have

$$
E_{1}=E_{1}^{\prime}
$$

where $\sigma^{\prime}=\sigma$
14.2 The capacitance moves along the $x$ direction which is perpendicular to the electric field of the capacitance.


In the frame $S$ ' where the charges are at rest.

$$
\begin{array}{ll}
E_{1}{ }^{\prime}=0 & B_{1}{ }^{\prime}=0 \\
E_{2}{ }^{\prime}=0 & B_{2}{ }^{\prime}=0 \\
E_{3}{ }^{\prime}=\frac{\sigma^{\prime}}{\varepsilon_{0}} & B_{3}{ }^{\prime}=0
\end{array}
$$

where

$$
\sigma=\gamma \sigma^{\prime}
$$

$$
\begin{aligned}
& E_{1}=E_{1}{ }^{\prime}=0 \\
& E_{2}=\gamma\left(E_{2}{ }^{\prime}+c \beta B_{3}{ }^{\prime}\right)=0 \\
& E_{3}=\gamma\left(-c \beta B_{2}{ }^{\prime}+E_{3}{ }^{\prime}\right)=\gamma E_{3}{ }^{\prime}=\gamma \frac{\sigma^{\prime}}{\varepsilon_{0}}=\frac{\sigma}{\varepsilon_{0}} \\
& B_{1}=B_{1}{ }^{\prime}=0 \\
& B_{2}=\gamma\left(B_{2}{ }^{\prime}-\frac{\beta}{c} E_{3}{ }^{\prime}\right)=\frac{-\beta E_{3}{ }^{\prime}}{c \sqrt{1-\beta^{2}}}=-\gamma \frac{v}{c^{2}} \frac{\sigma^{\prime}}{\varepsilon_{0}}=-\frac{v}{c^{2}} \frac{\sigma}{\varepsilon_{0}} \\
& B_{3}=\gamma\left(\frac{\beta}{c} E_{2}{ }^{\prime}+B_{3}{ }^{\prime}\right)=0
\end{aligned}
$$

## 15. Relativistic-covariant Lagrangian formalism

### 15.1 Lagrangian $L$ (simple case)

Proper time

$$
\left(d x_{\mu}^{\prime}\right)^{2}=a_{\mu \lambda} a_{\mu \sigma} d x_{\lambda} d x_{\sigma}=\delta_{\lambda \sigma} d x_{\lambda} d x_{\sigma}=\left(d x_{\mu}\right)^{2}
$$

We define the proper time as

$$
\begin{aligned}
& (d s)^{2}=c^{2}(d t)^{2}-\left(d x_{1}\right)^{2}-\left(d x_{2}\right)^{2}-\left(d x_{3}\right)^{2}=c^{2}\left(d t^{\prime}\right)^{2}-\left(d x_{1}{ }^{\prime}\right)^{2}-\left(d x_{2}{ }^{\prime}\right)^{2}-\left(d x_{3}{ }^{\prime}\right)^{2} \\
& (d s)^{2}=c^{2}(d t)^{2}\left\{1-\frac{1}{c^{2}}\left[\left(\frac{d x_{1}}{d t}\right)^{2}+\left(\frac{d x_{2}}{d t}\right)^{2}+\left(\frac{d x_{3}}{d t}\right)^{2}\right]\right\}=c^{2}(d t)^{2}\left(1-\frac{\mathbf{u}^{2}}{c^{2}}\right)
\end{aligned}
$$

or

$$
d \tau=\frac{d s}{c}=d t \sqrt{1-\frac{\mathbf{u}^{2}}{c^{2}}}
$$

where $\tau$ is a proper time and $\boldsymbol{u}$ is the velocity of the particle in the frame $S$.
The integral $\int_{a}^{b} d s$ taken between a given pair of world points has its maximum value if it is taken along the straight line joining two points.

$$
S=-\alpha \int_{a}^{b} d s=-\alpha c \int_{t_{a}}^{t_{b}} d t \sqrt{1-\frac{\mathbf{u}^{2}}{c^{2}}}=\int_{t_{a}}^{t_{b}} L d t
$$

where

$$
L=-\alpha c \sqrt{1-\frac{\mathbf{u}^{2}}{c^{2}}}
$$

Nonrelativistic case

$$
L=-\alpha c\left(1-\frac{\mathbf{u}^{2}}{c^{2}}\right)^{1 / 2}=-\alpha c\left(1-\frac{\mathbf{u}^{2}}{2 c^{2}}\right)=\frac{\alpha}{2 c} \mathbf{u}^{2}-\alpha c
$$

In the classical mechanics,

$$
\frac{\alpha}{2 c}=\frac{m}{2} \quad \text { or } \quad \alpha=m c
$$

Therefore the Lagrangian $L$ is given by

$$
L=-m c^{2}\left(1-\frac{\mathbf{u}^{2}}{c^{2}}\right)^{1 / 2}
$$

The momentum $\boldsymbol{p}$ is defined by

$$
\mathbf{p}=\frac{\partial L}{\partial \mathbf{u}}=\frac{m \mathbf{u}}{\sqrt{1-\frac{\mathbf{u}^{2}}{c^{2}}}}=m \mathbf{u} \gamma(\mathbf{u})=m \frac{d \mathbf{r}}{d \tau}=m \frac{d \mathbf{r}}{d t} \frac{d t}{d \tau}
$$

((Note))
This momentum coincides with the components of four-vector momentum $p_{\mu}$ defined by

$$
p_{\mu}=m \frac{d x_{\mu}}{d \tau}
$$

### 15.2 Hamiltonian

The Hamiltonian $H$ is defined by

$$
H=\mathbf{p} \cdot \mathbf{u}-L=\gamma(\mathbf{u}) m \mathbf{u}^{2}+m c^{2} \frac{1}{\gamma(\mathbf{u})}=\frac{\gamma(\mathbf{u})^{2} m \mathbf{u}^{2}+m c^{2}}{\gamma(\mathbf{u})}=\gamma(\mathbf{u}) m c^{2}=\frac{m c^{2}}{\sqrt{1-\frac{\mathbf{u}^{2}}{c^{2}}}}=E
$$

or

$$
E=\frac{m c^{2}}{\sqrt{1-\frac{\mathbf{u}^{2}}{c^{2}}}}
$$

We have

$$
\frac{E^{2}}{c^{2}}=\frac{m^{2} c^{2}}{1-\frac{\mathbf{u}^{2}}{c^{2}}}=\frac{m^{2} c^{2}\left(1-\frac{\mathbf{u}^{2}}{c^{2}}\right)+m^{2} \mathbf{u}^{2}}{1-\frac{\mathbf{u}^{2}}{c^{2}}}=m^{2} c^{2}+\mathbf{p}^{2}
$$

### 15.3 Lagrangian form in the presence of an electromagnetic field

The action function for a charge in an electromagnetic field

$$
S=\int_{a}^{b}\left(-m c d s+q A_{\mu} d x_{\mu}\right)
$$

where the second term is invariant under the Lorentz transformation.

$$
A_{\mu}=\left(\mathbf{A}, i \frac{1}{c} \phi\right), \quad \text { and } \quad d x_{\mu}=\left(d x_{1}, d x_{2}, d x_{3}, i c d t\right)
$$

Then we have

$$
S=\int_{a}^{b}\left(-m c d s+q A_{\mu} d x_{\mu}\right)=\int_{a}^{b}\left[-m c^{2} \sqrt{1-\frac{\mathbf{u}^{2}}{c^{2}}}+q(\mathbf{A} \cdot \mathbf{u}-\phi)\right] d t
$$

The integrand in the Lagrangian function of a charge $(q)$ in the electromagnetic field,

$$
\begin{aligned}
& L=-m c^{2} \sqrt{1-\frac{\mathbf{u}^{2}}{c^{2}}}+q(\mathbf{A} \cdot \mathbf{u}-\phi) \\
& \mathbf{p}=\frac{\partial L}{\partial \mathbf{u}}=\frac{m \mathbf{u}}{\sqrt{1-\frac{\mathbf{u}^{2}}{c^{2}}}}+q \mathbf{A}
\end{aligned}
$$

where

$$
A_{\mu}=\left(\mathbf{A}, i \frac{1}{C} \phi\right)
$$

The Hamiltonian $H$ is given by

$$
H=\mathbf{p} \cdot \mathbf{u}-L=\frac{m \mathbf{u}^{2}}{\sqrt{1-\frac{\mathbf{u}^{2}}{c^{2}}}}+e \mathbf{A} \cdot \mathbf{u}-\left(-m c^{2} \sqrt{1-\frac{\mathbf{u}^{2}}{c^{2}}}+q \mathbf{A} \cdot \mathbf{u}-q \phi\right)
$$

or

$$
\begin{aligned}
& H=\mathbf{p} \cdot \mathbf{u}-L=\frac{m \mathbf{u}^{2}}{\sqrt{1-\frac{\mathbf{u}^{2}}{c^{2}}}}+e \mathbf{A} \cdot \mathbf{u}-\left(-m c^{2} \sqrt{1-\frac{\mathbf{u}^{2}}{c^{2}}}+q \mathbf{A} \cdot \mathbf{u}-q \phi\right) \\
& H=\frac{m c^{2}}{\sqrt{1-\frac{\mathbf{u}^{2}}{c^{2}}}}+q \phi
\end{aligned}
$$

or

$$
\left(\frac{H-q \phi}{c}\right)^{2}=\frac{m^{2} c^{2}\left(1-\frac{\mathbf{u}^{2}}{c^{2}}\right)+m^{2} \mathbf{u}^{2}}{1-\frac{\mathbf{u}^{2}}{c^{2}}}=m^{2} c^{2}+(\mathbf{p}-q \mathbf{A})^{2}
$$

### 15.4 Expression for the Lagrangian in terms of 4-dimensional velocity

Here we use $\mathrm{d} \tau$ instead of $\mathrm{d} t$ in the expression of Lagrangian.

$$
d s=c d \tau
$$

$\eta_{\mu}$ is a four-dimensional velocity defined by

$$
\begin{aligned}
& \eta_{\mu}=\frac{d x_{\mu}}{d \tau}=\frac{d t}{d \tau} \frac{d x_{\mu}}{d t}=\left(\gamma(\mathbf{u}) u_{1}, \gamma(\mathbf{u}) u_{2}, \gamma(\mathbf{u}) u_{3}, i c \gamma(\mathbf{u})\right) \\
& A_{\mu} \eta_{\mu}=A_{1} \eta_{1}+A_{2} \eta_{2}+A_{3} \eta_{3}+A_{4} \eta_{4}=\gamma(\mathbf{u})(\mathbf{u} \cdot \mathbf{A}-\phi)
\end{aligned}
$$

since

$$
\begin{aligned}
& A_{\mu}=\left(\mathbf{A}, i \frac{1}{c} \phi\right), \quad \eta_{4}=\frac{d t}{d \tau} \frac{d x_{4}}{d t}=i c \frac{d t}{d \tau} \\
& S=\int_{a}^{b}\left(-m c d s+q A_{\mu} d x_{\mu}\right)=\int_{a}^{b}\left(-m c^{2}+q A_{\mu} \cdot \eta_{\mu}\right) d \tau
\end{aligned}
$$

$$
L=-m c^{2}+q A_{\mu} \eta_{\mu}
$$

### 15.5 Lagrangian and Hamiltonian in terms of the field tensor $\boldsymbol{F}_{\mu \nu}$

$$
F_{\mu \nu} F_{\mu \nu}=2\left(B_{1}^{2}+B_{2}^{2}+B_{3}^{2}\right)-\frac{2}{c^{2}}\left(E_{1}^{2}+E_{2}^{2}+E_{3}^{2}\right)
$$

This is invariant under the Lorentz transformation.
We may try the Lagrangian density

$$
L=-\frac{1}{4 \mu_{0}} F_{\mu \nu} F_{\mu \nu}+J_{\mu} A_{\mu}
$$

By regarding each component of $A_{\mu}$ as an independent field, we find that the Lagrange equation

$$
\frac{\partial L}{\partial A_{\mu}}=\frac{\partial}{\partial x_{v}}\left[\frac{\partial L}{\partial\left(\frac{\partial A_{\mu}}{\partial x_{v}}\right)}\right]
$$

is equivalent to

$$
\frac{\partial F_{\mu \nu}}{\partial x_{\mu}}=\mu_{0} J_{\mu} .
$$

The Hamiltonian density $H_{e m}$ for the free Maxwell field can be evaluated as follows.

$$
\begin{aligned}
& L_{e m}=-\frac{1}{4 \mu_{0}} F_{\mu \nu} F_{\mu \nu} \\
& H_{e m}=\frac{\partial L_{e m}}{\partial\left(\frac{\partial A_{\mu}}{\partial x_{4}}\right)} \frac{\partial A_{\mu}}{\partial x_{4}}-L_{e m}=-\frac{F_{4 \mu}}{\mu_{0}}\left(F_{4 \mu}+\frac{\partial A_{4}}{\partial x_{\mu}}\right)-\frac{1}{2 \mu_{0}}\left(\mathbf{B}^{2}-\frac{1}{c^{2}} \mathbf{E}^{2}\right)
\end{aligned}
$$

or

$$
H_{e m}=\frac{1}{2} \varepsilon_{0} \mathbf{E}^{2}+\frac{1}{2 \mu_{0}} \mathbf{B}^{2}-\varepsilon_{0} \mathbf{E} \cdot \nabla \phi
$$

$$
\int H_{e m} d \mathbf{r}=\frac{1}{2} \int\left(\varepsilon_{0} \mathbf{E}^{2}+\frac{1}{2 \mu_{0}} \mathbf{B}^{2}\right) d \mathbf{r}-\int \varepsilon_{0}(\mathbf{E} \cdot \nabla \phi) d \mathbf{r}=\frac{1}{2} \int\left(\varepsilon_{0} \mathbf{E}^{2}+\frac{1}{2 \mu_{0}} \mathbf{B}^{2}\right) d r
$$

((Note))

$$
\int(\mathbf{E} \cdot \nabla \phi) d \mathbf{r}=\int[\nabla \cdot(\mathbf{E} \phi)-\phi \nabla \cdot \mathbf{E}] d \mathbf{r}=\int \nabla \cdot(\mathbf{E} \phi) d \mathbf{r}=\int(\mathbf{E} \phi) \cdot d \mathbf{a}=0
$$

where $\mathbf{E} \phi$ vanishes sufficiently rapidly at infinity.

$$
\nabla \cdot \mathbf{E}=\frac{\rho}{\varepsilon_{0}}=0 \text { (in this case). }
$$

## 16. Relativistic form of Newton's law

### 16.1 Relativistic force

We define the force $\boldsymbol{F}$ and the kinetic energy $T$ as

$$
\begin{aligned}
& \mathbf{F}=\frac{d \mathbf{p}}{d t} \\
& \frac{d T}{d t}=A=\mathbf{F} \cdot \mathbf{u}
\end{aligned}
$$

where $\boldsymbol{u}$ is the velocity of the particle.

$$
\frac{d T}{d t}=\mathbf{u} \cdot \mathbf{F}=\mathbf{u} \cdot \frac{d \mathbf{p}}{d t}=\mathbf{u} \cdot \frac{d}{d t} \frac{m \mathbf{u}}{\sqrt{1-\frac{\mathbf{u}^{2}}{c^{2}}}}=m \mathbf{u} \cdot \frac{d \mathbf{u}}{d t} \frac{1}{\sqrt{1-\frac{\mathbf{u}^{2}}{c^{2}}}}+m \mathbf{u}^{2} \frac{d}{d t} \frac{1}{\sqrt{1-\frac{\mathbf{u}^{2}}{c^{2}}}}
$$

or

$$
A=\frac{d T}{d t}=m \mathbf{u} \cdot \frac{d \mathbf{u}}{d t} \frac{1}{\left(1-\frac{\mathbf{u}^{2}}{c^{2}}\right)^{1 / 2}}+m \mathbf{u}^{2} \frac{\mathbf{u} \cdot \frac{d \mathbf{u}}{d t}}{c^{2}\left(1-\frac{\mathbf{u}^{2}}{c^{2}}\right)^{3 / 2}}
$$

or

$$
\left.A=\frac{d T}{d t}=m \mathbf{u} \cdot \frac{d \mathbf{u}}{d t}\left[\frac{1-\frac{\mathbf{u}^{2}}{c^{2}}}{\left(1-\frac{\mathbf{u}^{2}}{c^{2}}\right)^{3 / 2}}+\frac{\frac{\mathbf{u}^{2}}{c^{2}}}{\left(1-\frac{\mathbf{u}^{2}}{c^{2}}\right)^{3 / 2}}\right]=\frac{m \mathbf{u} \cdot \frac{d \mathbf{u}}{d t}}{\left(1-\frac{\mathbf{u}^{2}}{c^{2}}\right)^{3 / 2}}=\frac{\frac{m}{2} \frac{d \mathbf{u}^{2}}{d t}}{\left(1-\frac{\mathbf{u}^{2}}{c^{2}}\right)^{3 / 2}}\right]
$$

$$
\begin{aligned}
& d T=\frac{m}{2} \frac{1}{\left(1-\frac{\mathbf{u}^{2}}{c^{2}}\right)^{3 / 2}} d \mathbf{u}^{2} \\
& T=\frac{m c^{2}}{\sqrt{1-\frac{\mathbf{u}^{2}}{c^{2}}}}+C
\end{aligned}
$$

where $C$ is a constant of integration. Since the kinetic energy may be taken as zero for $u=$ 0 .

### 16.2 Relativistic energy

Then we have $C==m c^{2}$.

$$
T=\frac{m c^{2}}{\sqrt{1-\frac{\mathbf{u}^{2}}{c^{2}}}}-m c^{2}
$$

It is convenient to introduce a quantity $E$ defined by

$$
E=T+m c^{2}=\frac{m c^{2}}{\sqrt{1-\frac{\mathbf{u}^{2}}{c^{2}}}}=\gamma(\mathbf{u}) m c^{2}
$$

$E$ is the energy of a free particle

## 17. Four-dimensional momentum

17.1. Definition

$$
\begin{aligned}
& d s=c d t \sqrt{1-\frac{\mathbf{u}^{2}}{c^{2}}}=\frac{c d t}{\gamma(\mathbf{u})} \\
& d \tau=\frac{d s}{c}=\frac{d t}{\gamma(\mathbf{u})}
\end{aligned}
$$

where $\tau$ is a proper time.

$$
\gamma(\mathbf{u})=\frac{1}{\sqrt{1-\frac{\mathbf{u}^{2}}{c^{2}}}}
$$

Here we define the four-dimensional momentum

$$
p=m \frac{d}{d \tau} x=m \gamma(\mathbf{u}) \frac{d x_{\mu}}{d t}=\left(m \gamma(\mathbf{u}) u_{1}, m \gamma(\mathbf{u}) u_{2}, m \gamma(\mathbf{u}) u_{3}, i m c \gamma(\mathbf{u})\right)=\left(\mathbf{p}, i \frac{E}{c}\right)
$$

This is exactly the same as the expressions of $\boldsymbol{p}$ obtained from the Lagrangian.

$$
E=m c^{2} \gamma(\mathbf{u})=\frac{m c^{2}}{\sqrt{1-\frac{\mathbf{u}^{2}}{c^{2}}}}
$$

### 17.2 Lorentz transformation

This momentum is clearly a four-vector since $d x_{\mu}$ is a Lorentz four-vector and $m$ and $\mathrm{d} \tau$ are Lorentz scalar. In fact, under the Lorentz transformation

$$
\begin{aligned}
& p_{\mu}^{\prime}=m \frac{d}{d \tau} x_{\mu}^{\prime}=m a_{\mu \nu} \frac{d}{d s} x_{v}=a_{\mu \nu} p_{v} \\
& \left(p_{\mu}^{\prime}\right)^{2}=a_{\mu \nu} a_{\mu \lambda} p_{v} p_{\lambda}=\delta_{v \lambda} p_{v} p_{\lambda}=\left(p_{\mu}\right)^{2}
\end{aligned}
$$

So this is invariant under the Lorentz transformation.

$$
\begin{array}{ll}
p=\left(\mathbf{p}, i \frac{E}{c}\right) & \\
p_{1}^{\prime}=\frac{p_{1}+i \beta p_{4}}{\sqrt{1-\beta^{2}}} & p_{1}^{\prime}=\frac{p_{1}-\frac{\beta}{c} E}{\sqrt{1-\beta^{2}}} \\
p_{2}^{\prime}=p_{2} & \text { or } \\
p_{3}^{\prime}=p_{3} & p_{2}^{\prime}=p_{2} \\
p_{4}^{\prime}=\frac{-i \beta p_{1}+p_{4}}{\sqrt{1-\beta^{2}}} & p_{3}^{\prime}=p_{3} \\
& E^{\prime}=\frac{E-\beta c p_{1}}{\sqrt{1-\beta^{2}}}
\end{array}
$$

and

$$
\begin{array}{ll}
p_{1}=\frac{p_{1}^{\prime}-i \beta p_{4}^{\prime}}{\sqrt{1-\beta^{2}}} & p_{1}=\frac{p_{1}^{\prime}+\frac{\beta}{c} E^{\prime}}{\sqrt{1-\beta^{2}}} \\
p_{2}=p_{2}^{\prime} & p_{2}=p_{2}^{\prime} \\
p_{3}=p_{3}^{\prime} & p_{3}=p_{3}^{\prime} \\
p_{4}=\frac{i \beta p_{1}^{\prime}+p_{4}^{\prime}}{\sqrt{1-\beta^{2}}} & E=\frac{E^{\prime}+\beta c p_{1}^{\prime}}{\sqrt{1-\beta^{2}}}
\end{array}
$$

18. Four-dimensional velocity (or proper 4-velocity)

$$
\begin{aligned}
& d s=c d t \sqrt{1-\frac{\mathbf{u}^{2}}{c^{2}}}=\frac{c d t}{\gamma(\mathbf{u})} \\
& d \tau=\frac{d s}{c}=\frac{d t}{\gamma(\mathbf{u})}
\end{aligned}
$$

where $\tau$ is a proper time. $\gamma(\mathbf{u})=\frac{1}{\sqrt{1-\frac{\mathbf{u}^{2}}{c^{2}}}}$

Here we define the four-dimensional velocity by

$$
\eta_{\mu}=\frac{d x_{\mu}}{d \tau}=\frac{d t}{d \tau} \frac{d x_{\mu}}{d t}=\left(\gamma(\mathbf{u}) u_{1}, \gamma(\mathbf{u}) u_{2}, \gamma(\mathbf{u}) u_{3}, i c \gamma(\mathbf{u})\right)
$$

where $u_{\mathrm{i}}$ is the 3-velocity

$$
\begin{aligned}
& u_{i}=\frac{d x_{i}}{d t} \quad(i=1,2,3) \\
& \eta_{\mu}^{\prime}=a_{\mu \nu} \eta_{v} \\
& \left(\begin{array}{l}
\eta_{1}{ }^{\prime} \\
\eta_{2}^{\prime} \\
\eta_{3}^{\prime} \\
\eta_{4}^{\prime}
\end{array}\right)=\left(\begin{array}{cccc}
\gamma & 0 & 0 & i \beta \gamma \\
0 & 1 & 0 & 0 \\
0 & 0 & 1 & 0 \\
-i \beta \gamma & 0 & 0 & \gamma
\end{array}\right)\left(\begin{array}{l}
\eta_{1} \\
\eta_{2} \\
\eta_{3} \\
\eta_{4}
\end{array}\right)
\end{aligned}
$$

or

$$
\eta_{\mu}=a_{v \mu} \eta_{v}{ }^{\prime}
$$

$$
\left(\begin{array}{l}
\eta_{1} \\
\eta_{2} \\
\eta_{3} \\
\eta_{4}
\end{array}\right)=\left(\begin{array}{cccc}
\gamma & 0 & 0 & -i \beta \gamma \\
0 & 1 & 0 & 0 \\
0 & 0 & 1 & 0 \\
i \beta \gamma & 0 & 0 & \gamma
\end{array}\right)\left(\begin{array}{l}
\eta_{1}{ }^{\prime} \\
\eta_{2}{ }^{\prime} \\
\eta_{3}{ }^{\prime} \\
\eta_{4}{ }^{\prime}
\end{array}\right)
$$

with

$$
\gamma=\frac{1}{\sqrt{1-\frac{v^{2}}{c^{2}}}}
$$

## 21. Force

$$
\begin{aligned}
& \frac{d p_{1}}{d t}=\frac{d t}{d t^{\prime}} \frac{d p_{1}}{d t}=\frac{1}{\gamma\left(1+\frac{\beta}{c} u_{1}{ }^{\prime}\right)}\left[\frac{d p_{1}^{\prime}}{d t^{\prime}}+m c \gamma \beta \frac{d}{d t^{\prime}} \gamma\left(\mathbf{u}^{\prime}\right)\right] \\
& \frac{d p_{2}}{d t}=\frac{d t^{\prime}}{d t} \frac{d p_{2}}{d t^{\prime}}=\frac{d p_{2}^{\prime}}{d t^{\prime}} \frac{1}{\gamma\left(1+\frac{\beta}{c} u_{1}^{\prime}\right)} \\
& \frac{d p_{3}}{d t}=\frac{d t^{\prime}}{d t} \frac{d p_{3}}{d t^{\prime}}=\frac{d p_{3}^{\prime}}{d t^{\prime}} \frac{1}{\gamma\left(1+\frac{\beta}{c} u_{1}^{\prime}\right)}
\end{aligned}
$$

When $u_{1}{ }^{\prime}=0$ and $d u_{1}{ }^{\prime} / d t^{\prime}=0$

$$
\begin{aligned}
& \frac{d p_{1}}{d t}=\frac{d t^{\prime}}{d t} \frac{d p_{1}}{d t^{\prime}}=\frac{1}{\gamma\left(1+\frac{\beta}{c} u_{1}^{\prime}\right)}\left[\frac{d p_{1}^{\prime}}{d t^{\prime}}+m c \gamma \beta \frac{d}{d t^{\prime}} \gamma\left(\mathbf{u}^{\prime}\right)\right] \\
& \frac{d p_{2}}{d t}=\frac{1}{\gamma} \frac{d p_{2}^{\prime}}{d t^{\prime}} \\
& \frac{d p_{3}}{d t}=\frac{1}{\gamma} \frac{d p_{3}^{\prime}}{d t^{\prime}}
\end{aligned}
$$

## 22. Minkowski force

We define the Minkowski force as

$$
K_{\mu}=\frac{d p_{\mu}}{d \tau}
$$

This is a 4-dimesional vector. The spatial components of $K_{\mu}$ are related to the ordinary force by

$$
\mathbf{K}=\frac{d \mathbf{p}}{d \tau}=\frac{d t}{d \tau} \frac{d \mathbf{p}}{d t}=\frac{1}{\sqrt{1-\mathbf{u}^{2}}} \frac{d \mathbf{p}}{d t}=\frac{1}{\sqrt{1-\mathbf{u}^{2}}} \mathbf{F}
$$

The $4^{\text {th }}$ componenet

$$
K_{4}=\frac{d p_{4}}{d \tau}=\frac{d t}{d \tau} \frac{d p_{4}}{d t}=\frac{1}{c} \frac{d E}{d t}
$$

## 23 Lorentz force in the relativistic mechanics

23.1

$$
\mathbf{F}=\frac{d \mathbf{p}}{d t}=q[\mathbf{E}+(\mathbf{v} \times \mathbf{B})]
$$

holds in an arbitrary frame $S$. This expression is the correct relativistic form for Newton's second law. The momentum form is more fundamental.

The four-dimensional momentum is given by

$$
\begin{aligned}
& \mathbf{p}=m \frac{\mathbf{v}}{\sqrt{1-\frac{\mathbf{v}^{2}}{c^{2}}}} \\
& E_{k i n}=\frac{m c^{2}}{\sqrt{1-\frac{\mathbf{v}^{2}}{c^{2}}}} \\
& \frac{E^{2}}{c^{2}}=\frac{m^{2} c^{2}}{1-\frac{\mathbf{v}^{2}}{c^{2}}}=\frac{m^{2} c^{2}\left(1-\frac{\mathbf{v}^{2}}{c^{2}}\right)+m^{2} \mathbf{v}^{2}}{1-\frac{\mathbf{v}^{2}}{c^{2}}}=m^{2} c^{2}+\mathbf{p}^{2}
\end{aligned}
$$

or

$$
E_{k i n}=c\left(m^{2} c^{2}+\mathbf{p}^{2}\right)^{1 / 2}
$$

The final form of the equation of motion is given by

$$
\begin{align*}
& \frac{d}{d t} \mathbf{p}=q[\mathbf{E}+(\mathbf{v} \times \mathbf{B})]  \tag{1}\\
& \mathbf{p}=\frac{\mathbf{v}}{\sqrt{1-\frac{\mathbf{v}^{2}}{c^{2}}}}, \\
& \frac{d E_{\text {kin }}}{d t}=\mathbf{F} \cdot \mathbf{v}=q(\mathbf{v} \cdot \mathbf{E})
\end{align*}
$$

where

$$
\begin{equation*}
E_{k i n}=\frac{m c^{2}}{\sqrt{1-\frac{\mathbf{v}^{2}}{c^{2}}}}=c \sqrt{m^{2} c^{2}+\mathbf{p}^{2}} \tag{2}
\end{equation*}
$$

## 24. Cyclotron motion: a particle in a uniform magnetic field along the $\mathbf{z}$ axis.

 We now consider the case of $\boldsymbol{E}=0$.$$
\frac{d}{d t} E_{k i n}=0
$$

Thus we have $\gamma(\mathbf{v})=\frac{1}{\sqrt{1-\frac{\mathbf{v}^{2}}{c^{2}}}}=\frac{E_{\text {kin }}}{m c^{2}}=$ constant
The momentum:

$$
\mathbf{p}=\frac{E_{k i n} \mathbf{v}}{c^{2}}
$$

The equation of motion

$$
\frac{d}{d t} \mathbf{v}=\frac{c^{2}}{E_{\text {kin }}} q(\mathbf{v} \times \mathbf{B})
$$

or

$$
\begin{aligned}
& \dot{v}_{x}=\frac{c^{2} q B}{E_{k i n}} v_{y} \\
& \dot{v}_{y}=-\frac{c^{2} q B}{E_{k i n}} v_{x} \\
& \dot{v}_{z}=0
\end{aligned}
$$

We use the complex plane for the solution.

$$
\frac{d}{d t}\left(v_{x}+i v_{y}\right)=-\frac{i c^{2} q B}{E_{k i n}}\left(v_{x}+i v_{y}\right)
$$

or

$$
\left(v_{x}+i v_{y}\right)=\left(v_{x}^{0}+i v_{y}{ }^{0}\right) \exp \left[-\frac{i c^{2} q B t}{E_{k i n}}\right]=v \exp [-i(\omega t+\alpha)]
$$

where

$$
\begin{aligned}
& \omega=\frac{c^{2} q B}{E_{k i n}} \\
& v_{x}^{0}+i v_{y}^{0}=v e^{-i \alpha}
\end{aligned}
$$

Then we have

$$
\begin{aligned}
& v_{x}=\frac{d x}{d t}=v \cos (\omega t+\alpha) \\
& v_{x}=\frac{d y}{d t}=-v \sin (\omega t+\alpha)
\end{aligned}
$$

or

$$
\begin{aligned}
& v_{x}^{2}+v_{y}^{2}=v^{2}=\text { cons } \tan t \\
& x=\frac{v}{\omega} \sin (\omega t+\alpha)+x_{1} \\
& y=\frac{v}{\omega} \cos (\omega t+\alpha)+y_{1}
\end{aligned}
$$

This equation describes a cyclotron motion (circular motion with radius $R$ ).

$$
R=\frac{v}{\omega}=\frac{v E_{k i n}}{c^{2} q B}=\frac{p}{q B}
$$

where w is the angular frequency,

$$
\omega=\frac{c^{2} q B}{E_{k i n}}
$$

or

$$
B R=\frac{1}{q} p
$$

The radius has a maximum when $\frac{v}{c}=\frac{1}{\sqrt{2}}$
In summary

$$
\begin{aligned}
& x=\frac{\sqrt{p_{0 x}{ }^{2}+p_{0 y}{ }^{2}}}{q B} \sin \left(\frac{c^{2} q B}{E_{\text {kin }}} t+\alpha\right) \\
& y=\frac{\sqrt{p_{0 x}{ }^{2}+p_{0 y}{ }^{2}}}{q B} \cos \left(\frac{c^{2} q B}{E_{\text {kin }}} t+\alpha\right)
\end{aligned}
$$

25. The motion of the particle under an electric field $(E=-\nabla \phi)$

$$
\frac{d}{d t} E_{k i n}=q(\mathbf{v} \cdot \mathbf{E})=-q v \cdot \nabla \phi=-q \frac{d}{d t} \phi
$$

or

$$
\frac{d}{d t}\left(E_{k i n}+q \phi\right)=0
$$

or

$$
E_{k i n}+q \phi=\text { constant }
$$

We now consider the capacitance consisting of two parallel planes. Suppose that the particle with charge $q$ on the one plate moves to the other plate. The initial velocity is equal to zero. What is the velocity of the particle arriving at the other plate?

$$
E_{k i n}+q \phi_{2}=m c^{2}+q \phi_{1}
$$

When $\phi=\phi_{1}-\phi_{2}$,

$$
\frac{1}{1-\frac{\mathbf{v}^{2}}{c^{2}}}=\left(1+\frac{q \phi}{m c^{2}}\right)^{2}
$$

or

$$
v=c\left[1-\frac{1}{\left(1+\frac{q \phi}{m c^{2}}\right)^{2}}\right]^{1 / 2}
$$

## 26. Equation of motion under a constant electric field

We assume that $\boldsymbol{E}$ is along the $y$ axis. The initial momentum $\boldsymbol{p}_{0}$ is in the $(x, y)$ plane. The particle is at the origin at $t=0$.

$$
\begin{aligned}
& \frac{d}{d t} \mathbf{p}=q \mathbf{E} \\
& \mathbf{p}=\mathbf{p}_{0}+q \mathbf{E} t
\end{aligned}
$$

or

$$
\begin{aligned}
& \mathbf{p}=\left(p_{0 x}, q E t+p_{0 y}, 0\right) \\
& E_{k i n}=c\left(m^{2} c^{2}+\mathbf{p}^{2}\right)^{1 / 2}=\left[m^{2} c^{4}+c^{2}\left(p_{0 x}{ }^{2}+p_{0 y}{ }^{2}\right)+c^{2}\left(q^{2} E^{2} t^{2}+2 p_{0 y} q E t\right)\right]^{1 / 2}
\end{aligned}
$$

or

$$
E_{k i n}=\left[\left(E_{k i n}{ }^{0}\right)^{2}+c^{2}\left(q^{2} E^{2} t^{2}+2 p_{0 y} q E t\right)\right]^{1 / 2}
$$

where $E_{\text {kin }}{ }^{0}$ is the kinetic energy at the beginning of the motion $(t=0)$.

$$
\begin{aligned}
& E_{k i n}{ }^{0}=\sqrt{m^{2} c^{4}+c^{2}\left(p_{0 x}{ }^{2}+p_{0 y}{ }^{2}\right)} \\
& \mathbf{p}=\frac{m \mathbf{v}}{\sqrt{1-\frac{\mathbf{v}^{2}}{c^{2}}}}=m \mathbf{v} \frac{E_{k i n}}{m c^{2}}=\frac{E_{k i n} \mathbf{v}}{c^{2}}
\end{aligned}
$$

Thus we have
$\mathbf{v}=\frac{c^{2}}{E_{\text {kin }}} \mathbf{p}=\frac{c^{2}}{E_{\text {kin }}}\left(p_{0 x}, q E t+p_{0 y}, 0\right)=\frac{c^{2}}{\left[\left(E_{\text {kin }}{ }^{0}\right)^{2}+c^{2}\left(q^{2} E^{2} t^{2}+2 p_{0 y} q E t\right)\right]^{1 / 2}}\left(p_{0 x}, q E t+p_{0 y}, 0\right)$
or

$$
\begin{aligned}
& \frac{d x}{d t}=\frac{c^{2} p_{0 x}}{\left[\left(E_{\text {kin }}{ }^{0}\right)^{2}+c^{2}\left(q^{2} E^{2} t^{2}+2 p_{0 y} q E t\right)\right]^{1 / 2}} \\
& \frac{d y}{d t}=\frac{c^{2}\left(p_{0 y}+q E t\right)}{\left[\left(E_{k i n}^{0}\right)^{2}+c^{2}\left(q^{2} E^{2} t^{2}+2 p_{0 y} q E t\right)\right]^{1 / 2}} \\
& \frac{d z}{d t}=0
\end{aligned}
$$

Solving these differential equations (we use the Mathematica),

$$
y=\frac{1}{q E}\left[\sqrt{\left(E_{k i n}{ }^{0}\right)^{2}+c^{2} q^{2} E^{2} t^{2}+2 p_{0 y} c^{2} E q t}-E_{k i n}{ }^{0}\right]
$$

or

$$
\begin{aligned}
& y=\frac{c}{q E}\left[\sqrt{\left(p_{0 y}+q E t\right)^{2}+m^{2} c^{2}+p_{0 x}{ }^{2}}-\frac{E_{\text {kin }}{ }^{0}}{c}\right] \\
& x=\frac{c p_{0 x}}{q E} \ln \frac{p_{0 y}+q E t+\sqrt{\left(p_{0 y}+q E t\right)^{2}+m^{2} c^{2}+p_{0 x}{ }^{2}}}{p_{0 y}+\frac{E_{k i n}{ }^{0}}{c}}
\end{aligned}
$$

$$
z=0
$$

We now consider the special case when $p_{0 y}=0$.

$$
\frac{q E y}{c}=\sqrt{(q E t)^{2}+m^{2} c^{2}+p_{0 x}{ }^{2}}-\sqrt{m^{2} c^{2}+p_{0 x}{ }^{2}}
$$

$$
x=\frac{c p_{0 x}}{q E} \ln \frac{q E t+\sqrt{(q E t)^{2}+m^{2} c^{2}+p_{0 x}{ }^{2}}}{\sqrt{m^{2} c^{2}+p_{0 x}{ }^{2}}}
$$

or

$$
\begin{aligned}
& \exp \left(\frac{q E x}{c p_{0 x}}\right)=\frac{q E t+\sqrt{(q E t)^{2}+m^{2} c^{2}+p_{0 x}^{2}}}{\sqrt{m^{2} c^{2}+{p_{0 x}{ }^{2}}^{2}}} \\
& \exp \left(-\frac{q E x}{c p_{0 y}}\right)=\frac{\sqrt{m^{2} c^{2}+p_{0 x}^{2}}}{q E t+\sqrt{(q E t)^{2}+m^{2} c^{2}+p_{0 x}^{2}}}=\frac{-q E t+\sqrt{(q E t)^{2}+m^{2} c^{2}+p_{0 x}^{2}}}{\sqrt{m^{2} c^{2}+{p_{0 x}{ }^{2}}^{2}}} \\
& \cosh \left(\frac{q E x}{c p_{0 x}}\right)=\frac{\sqrt{(q E t)^{2}+m^{2} c^{2}+p_{0 x}^{2}}}{\sqrt{m^{2} c^{2}+p_{0 x}^{2}}} \\
& \frac{q E y}{c}+\sqrt{m^{2} c^{2}+{p_{0 x}}^{2}}=\sqrt{(q E t)^{2}+m^{2} c^{2}+{p_{0 x}}^{2}}=\sqrt{m^{2} c^{2}+p_{0 x}^{2}} \cosh \left(\frac{q E x}{c p_{0 x}}\right)
\end{aligned}
$$

or

$$
y=\frac{E_{\text {kin }}{ }^{0}}{q E}\left[\cosh \left(\frac{q E x}{c p_{0 x}}\right)-1\right]=\frac{E_{\text {kin }}{ }^{0}}{q E}\left[\cosh \left(\frac{q E x}{c p_{0 x}}\right)-1\right]=\frac{c}{q E} \sqrt{m^{2} c^{2}+{p_{0 x}}^{2}}\left[\cosh \left(\frac{q E x}{c p_{0 x}}\right)-1\right]
$$

Thus in a uniform electric field, a charge $q$ moves along a catenary curve.
((Mathematica-13))

## Zimmerman

2D motion of a relativistic particle in a uniform
electric field

$$
\begin{aligned}
& Y 1=\left\{\frac{p x 0}{m}, \frac{q E 0 t}{m}, 0\right\} \\
& \left\{\frac{p \times 0}{m}, \frac{E 0 q t}{m}, 0\right\}
\end{aligned}
$$

$$
\mathrm{eq} 1=\frac{\xi^{2}}{1-\frac{\xi^{2}}{c^{2}}}=\mathrm{Y} 1 . \mathrm{Y} 1
$$

$\frac{\xi^{2}}{1-\frac{\xi^{2}}{c^{2}}}=\frac{p x \Theta^{2}}{m^{2}}+\frac{E 0^{2} q^{2} t^{2}}{m^{2}}$

$$
\text { eq2=Solve }[\text { eq1, } \xi] / / \text { Simplify }
$$

$$
\left\{\left\{\xi \rightarrow-\frac{\sqrt{-\mathrm{c}^{2}\left(\mathrm{p} \times 0^{2}+\mathrm{E} 0^{2} \mathrm{q}^{2} \mathrm{t}^{2}\right)}}{\sqrt{-\mathrm{c}^{2} \mathrm{~m}^{2}-\mathrm{px} 0^{2}-} \mathrm{EO}^{2} \mathrm{a}^{2} \mathrm{q}^{2} \mathrm{t}^{2}}\right\},\left\{\xi \rightarrow \frac{\sqrt{-\mathrm{c}^{2}\left(\mathrm{px} 0^{2}+\mathrm{E} 0^{2} \mathrm{q}^{2} \mathrm{t}^{2}\right)}}{\sqrt{-\mathrm{c}^{2} \mathrm{~m}^{2}-\mathrm{px} 0^{2}-E \theta^{2} \mathrm{q}^{2} \mathrm{t}^{2}}}\right\}\right\}
$$

$$
\begin{aligned}
& e q 3=\sqrt{1-\frac{\xi^{2}}{c^{2}}} / . e q 2[[2]] / / \text { Simplify } \\
& \sqrt{\frac{c^{2} m^{2}}{c^{2} m^{2}+p \times 0^{2}+E 0^{2} q^{2} t^{2}}} \\
& \mathrm{~V}=\left\{\mathrm{x}^{\prime}[\mathrm{t}], \mathrm{y}^{\prime}[\mathrm{t}], \mathrm{z}^{\prime}[\mathrm{t}]\right\} \\
& \left\{\mathrm{X}^{\prime}[\mathrm{t}], \mathrm{y}^{\prime}[\mathrm{t}], \mathrm{z}^{\prime}[\mathrm{t}]\right\} \\
& \text { eq4=eq3 Y1 } \\
& \left\{\frac{p x 0 \sqrt{\frac{c^{2} m^{2}}{c^{2} m^{2}+p x 0^{2}+E 0^{2} q^{2} t^{2}}}}{m}, \frac{E 0 q t \sqrt{\frac{c^{2} m^{2}}{c^{2} m^{2}+p x 0^{2}+E 0^{2} q^{2} t^{2}}}}{m}, 0\right\} \\
& \text { eq5=Table[V[[i]]==eq4[[i]],\{i,1,3\}] } \\
& \left\{x^{\prime}[t]=\frac{p x 0 \sqrt{\frac{c^{2} m^{2}}{c^{2} m^{2}+p x 0^{2}+E 0^{2} q^{2} t^{2}}}}{m},\right. \\
& \left.y^{\prime}[t]=\frac{E 0 q t \sqrt{\frac{c^{2} m^{2}}{c^{2} m^{2}+p x 0^{2}+E 0^{2} q^{2} t^{2}}}}{m}, z^{\prime}[t]=0\right\} \\
& \text { eq6=DSolve }[\{\text { eq5, }\{x[0]==0, y[0]=0, z[0]=0\}\},\{x[t], y[t], z[t \\
& \text { ] \}, t] / /Simplify } \\
& \left\{\left\{x [ t ] \rightarrow \frac { 1 } { E 0 m q } \left(-p x 0 \sqrt{\frac{c^{2} m^{2}}{c^{2} m^{2}+p x 0^{2}}} \sqrt{c^{2} m^{2}+p x 0^{2}} \log \left[2 \sqrt{c^{2} m^{2}+p x 0^{2}}\right]+\right.\right.\right. \\
& p x 0 \sqrt{\frac{c^{2} m^{2}}{c^{2} m^{2}+p x 0^{2}+E 0^{2} q^{2} t^{2}}} \sqrt{c^{2} m^{2}+p x 0^{2}+E 0^{2} q^{2} t^{2}} \\
& \left.\log \left[2\left(E 0 q t+\sqrt{c^{2} m^{2}+p x 0^{2}+E 0^{2} q^{2} t^{2}}\right)\right]\right), \\
& \left.\left.y[t] \rightarrow \frac{c^{2} m\left(\sqrt{\frac{c^{2} m^{2}}{c^{2} m^{2}+p \times 0^{2}}}-\sqrt{\frac{c^{2} m^{2}}{c^{2} m^{2}+p \times 0^{2}+E^{2} q^{2} t^{2}}}\right)}{E 0 \sqrt{\frac{c^{2} m^{2}}{c^{2} m^{2}+p x 0^{2}}} q \sqrt{\frac{c^{2} m^{2}}{c^{2} m^{2}+p x 0^{2}+E^{2} q^{2} t^{2}}}}, z[t] \rightarrow 0\right\}\right\} \\
& \text { rule1 }=\{\mathrm{m} \rightarrow 1, \mathrm{q} \rightarrow 1, \mathrm{E} 0 \rightarrow 0.1, \mathrm{px} 0 \rightarrow 0.1, \mathrm{c} \rightarrow 1\} \\
& \{\mathrm{m} \rightarrow 1, \mathrm{q} \rightarrow 1, \mathrm{E} 0 \rightarrow 0.1, \mathrm{px} 0 \rightarrow 0.1, \mathrm{c} \rightarrow 1\} \\
& x\left[t \_\right]=x[t] / . e q 6[[1,1]] / . r u l e 1 / / S i m p l i f y \\
& -0.698122+ \\
& \text { 10. } \sqrt{1.01+0.01 t^{2}} \sqrt{\frac{1}{101 .+1 . t^{2}}} \log \left[2\left(0.1 t+\sqrt{1.01+0.01 t^{2}}\right)\right] \\
& y[t]=y[t] / \text {.eq6[ }[1,2]] / . r u l e 1 / / S i m p l i f y \\
& \text { 1. }-10.0499 \sqrt{\frac{1}{101 .+1 . t^{2}}} \\
& \int \frac{1}{101 \cdot+1 . t^{2}} \\
& z\left[t \_\right]=z[t] / \text {.eq6 [ [1, 3] ]/.rule1//Simplify }
\end{aligned}
$$

pll=Plot[\{x[t],y[t]\},\{t,0,30\},PlotStyle $\rightarrow$ Table [Hue [0.5 i],\{i,0,1\}], Prolog $\rightarrow$ AbsoluteThickness[2], PlotPoints $\rightarrow 50$, Background $\rightarrow$ GrayLevel [0.7]]

-Graphics-
Nonrelativistic motion
xnon $\left[\mathrm{t}_{-}\right]=\frac{\mathrm{px0t}}{\mathrm{~m}} /$. rule1
0.1 t
ynon $[t]=\frac{q E 0 t^{2}}{2 m} /$ rule1
$0.05 \mathrm{t}^{2}$
pl2=Plot[\{xnon[t],ynon[t]\},\{t, 0,30\},PlotStyle $\rightarrow$ Table [Hu e[0.3 i], $\{i, 1,2\}]$, Prolog $\rightarrow$ AbsoluteThickness[1.6], PlotPoints $\rightarrow 50$, Background $\rightarrow$ GrayLevel [0.7]]

-Graphics-
Show [pl1, pl2]

-Graphics-

## 28. A particle in a uniform electric field and a magnetic field

Let the electric field $\boldsymbol{E}$ be parallel to the $y$ axis and the magnetic field $\boldsymbol{B}$ parallel to the $z$ axis. At $t=0$ the particle is at the point $(0,0,0)$ and has a momentum $\boldsymbol{p}_{0}$.

Lorentz invariant:

$$
\begin{equation*}
\frac{d}{d t} \mathbf{p}=\mathbf{F}=q[\mathbf{E}+(\mathbf{v} \times \mathbf{B})] \tag{1}
\end{equation*}
$$

According to the Lorentz invariance, we have

$$
\begin{aligned}
& \mathbf{B}^{2}-\frac{1}{c^{2}} \mathbf{E}^{2}=\mathbf{B}^{\prime 2}-\frac{1}{c^{2}} \mathbf{E}^{\prime 2} \\
& \mathbf{E} \cdot \mathbf{B}=\mathbf{E}^{\prime} \cdot \mathbf{B}^{\prime}
\end{aligned}
$$

Since $\mathbf{E} \cdot \mathbf{B}=0$, we have $\mathbf{E}^{\prime} \cdot \mathbf{B}^{\prime}=0$

We assume a frame that $\mathbf{B}^{\prime}=0$.
In this case, we have

$$
\mathbf{B}^{2}-\frac{1}{c^{2}} \mathbf{E}^{2}=-\frac{1}{c^{2}} \mathbf{E}^{\prime 2}<0
$$

or

$$
\mathbf{B}^{2}<\frac{1}{c^{2}} \mathbf{E}^{2}
$$

or

$$
B_{3}<\frac{1}{C} E_{2}
$$

This is a condition for $\boldsymbol{E}$ and $\boldsymbol{B}$. Using the Lorentz transformation, we have

$$
\begin{array}{ll}
E_{1}^{\prime}=E_{1}=0 & B_{1}{ }^{\prime}=B_{1}=0 \\
E_{2}{ }^{\prime}=\gamma\left(E_{2}-c \beta B_{3}\right) & B_{2}{ }^{\prime}=\gamma\left(B_{2}+\frac{\beta}{c} E_{3}\right)=0 \\
E_{3}{ }^{\prime}=\gamma\left(E_{3}+c \beta B_{2}\right)=0 & B_{3}{ }^{\prime}=\gamma\left(B_{3}-\frac{\beta}{c} E_{2}\right)=0
\end{array}
$$

We choose $B_{3}{ }^{\prime}=0$

$$
B_{3}^{\prime}=B_{3}-\frac{\beta}{c} E_{2}=0
$$

or

$$
v=c^{2} \frac{B_{3}}{E_{2}}<c
$$

In this case,

$$
\begin{aligned}
& \mathbf{B}^{\prime}=0 \\
& E_{2}^{\prime}=\gamma\left(E_{2}-c \beta B_{3}\right)=\frac{1}{\gamma} E_{2}=\frac{E}{\gamma}=E^{\prime} \\
& E_{1}^{\prime}=0 \\
& E_{3}^{\prime}=0
\end{aligned}
$$

The frame $S^{\prime}$ move relative to the frame $S$ with a velocity $v$ along the $x$ axis. We know the equation of motion for the particle in a uniform electric field $\mathbf{E}^{\prime}$ along the $y$ axis.

$$
\begin{aligned}
& x^{\prime}=\frac{c p_{0 x^{\prime}}}{q E^{\prime}} \ln \frac{p_{0 y}{ }^{\prime}+q E^{\prime} t^{\prime}+\sqrt{\left(p_{0 y}{ }^{\prime}+q E^{\prime} t^{\prime}\right)^{2}+m^{2} c^{2}+p_{0 x}{ }^{\prime 2}}}{p_{0 y}{ }^{\prime}+\frac{E_{k i n}{ }^{0}{ }^{0}}{c}} \\
& y^{\prime}=\frac{c}{q E^{\prime}}\left[\sqrt{\left(p_{0 y}{ }^{\prime}+q E^{\prime} t^{\prime}\right)^{2}+m^{2} c^{2}+p_{0 x}{ }^{\prime 2}}-\frac{E_{k i n}{ }^{0}{ }^{\prime}}{c}\right] \\
& E_{\text {kin }}{ }^{10}=c\left(m^{2} c^{2}+p_{0 x}{ }^{12}+p_{0 y}{ }^{\prime 2}\right)^{1 / 2}
\end{aligned}
$$

with $v=c^{2} \frac{B_{3}}{E_{2}}<c$
The Lorentz transformation between $p_{\mu}{ }^{0}=\left(\mathbf{p}_{0}, i \frac{E_{\text {kin }}{ }^{0}}{c}\right)$ and $p_{\mu}{ }^{0}=\left(\mathbf{p}_{0}{ }^{\prime}, i \frac{E_{\text {kin }}{ }^{0}{ }^{1}}{c}\right)$ is given by

$$
\begin{aligned}
& p_{01}{ }^{\prime}=\gamma\left(p_{01}-\frac{\beta}{c} E_{k i n}{ }^{0}\right) \\
& p_{02}{ }^{\prime}=p_{02} \\
& p_{03}{ }^{\prime}=p_{03} \\
& E_{k i n}{ }^{\prime}{ }^{\prime}=\gamma\left(E_{\text {kin }}{ }^{0}-\beta c p_{01}\right)
\end{aligned}
$$

The required equations of motion for the particle in the frame S is obtained using the Lorentz transformation.

$$
\begin{array}{ll}
x_{1}=\gamma\left(x_{1}^{\prime}-i \beta x_{4}{ }^{\prime}\right) & x=\gamma\left(x^{\prime}+v t^{\prime}\right) \\
x_{2}=x_{2}^{\prime} & y=y^{\prime} \\
x_{3}=x_{3}^{\prime} & z=z^{\prime} \\
x_{4}=\gamma\left(i \beta x_{1}^{\prime}+x_{4}^{\prime}\right) & t=\gamma\left(\frac{\beta}{c} x^{\prime}+t^{\prime}\right)
\end{array}
$$

We assume a frame $S^{\prime}$ that $\mathbf{E}^{\prime}=0$.
In this case, we have

$$
\mathbf{B}^{2}-\frac{1}{c^{2}} \mathbf{E}^{2}=\mathbf{B}^{\prime 2}>0
$$

or

$$
\mathbf{B}^{2}>\frac{1}{c^{2}} \mathbf{E}^{2}
$$

or

$$
B_{3}>\frac{1}{c} E_{2}
$$

This is a condition for $\boldsymbol{E}$ and $\boldsymbol{B}$. Using the Lorentz transformation, we have

$$
\begin{array}{ll}
E_{1}^{\prime}=E_{1}=0 & B_{1}{ }^{\prime}=B_{1}=0 \\
E_{2}^{\prime}=\gamma\left(E_{2}-c \beta B_{3}\right) & B_{2}^{\prime}=\gamma\left(B_{2}+\frac{\beta}{c} E_{3}\right)=0 \\
E_{3}^{\prime}=\gamma\left(E_{3}+c \beta B_{2}\right)=0 & B_{3}^{\prime}=\gamma\left(B_{3}-\frac{\beta}{c} E_{2}\right)
\end{array}
$$

We choose $E_{2}{ }^{\prime}=0$

$$
E_{2}^{\prime}=\gamma\left(E_{2}-c \beta B_{3}\right)=0
$$

or

$$
v=\frac{E_{2}}{B_{3}}
$$

In this case,

$$
\begin{aligned}
& \mathbf{E}^{\prime}=0 \\
& B_{3}{ }^{\prime}=\gamma\left(B_{3}-\frac{\beta}{c} v B_{3}\right)=\frac{1}{\gamma} B_{3} \\
& B_{1}^{\prime}=0 \\
& B_{2}{ }^{\prime}=0
\end{aligned}
$$

The frame $S$ ' move relative to the frame $S$ with a velocity $v\left(=E_{2} / B_{3}<c\right)$ along the $x$ axis. We know the equation of motion for the particle in a uniform electric field $\mathbf{B}^{\prime}$ along the $z^{\prime}$ axis.

$$
\begin{aligned}
& x^{\prime}=\frac{\sqrt{p_{0 x}{ }^{\prime 2}+p_{0 y}{ }^{\prime 2}}}{q B^{\prime}} \sin \left(\frac{c^{2} q B^{\prime}}{E_{\text {kin }}{ }^{\prime}} t^{\prime}+\alpha^{\prime}\right) \\
& y^{\prime}=\frac{\sqrt{p_{0 x}{ }^{\prime 2}+p_{0 y}{ }^{\prime 2}}}{q B^{\prime}} \cos \left(\frac{c^{2} q B^{\prime}}{E_{k i n}{ }^{\prime}} t^{\prime}+\alpha^{\prime}\right)
\end{aligned}
$$

with $v=E_{2} / B_{3}<c$.
The Lorentz transformation between $p_{\mu}{ }^{0}=\left(\mathbf{p}_{0}, i \frac{E_{\text {kin }}{ }^{0}}{c}\right)$ and $p_{\mu}{ }^{0}=\left(\mathbf{p}_{0}{ }^{\prime}, i \frac{E_{\text {kin }}{ }^{0}{ }^{1}}{C}\right)$ is given by

$$
\begin{aligned}
& p_{01}{ }^{\prime}=\gamma\left(p_{01}-\frac{\beta}{c} E_{k i n}{ }^{0}\right) \\
& p_{02}{ }^{\prime}=p_{02} \\
& p_{03}{ }^{\prime}=p_{03} \\
& E_{\text {kin }}{ }^{0}{ }^{\prime}=\gamma\left(E_{\text {kin }}{ }^{0}-\beta c p_{01}\right)
\end{aligned}
$$

The required equations of motion for the particle in the frame S is obtained using the Lorentz transformation.

$$
\begin{array}{ll}
x_{1}=\gamma\left(x_{1}^{\prime}-i \beta x_{4}{ }^{\prime}\right) & x=\gamma\left(x^{\prime}+v t^{\prime}\right) \\
x_{2}=x_{2}{ }^{\prime} & y=y^{\prime} \\
x_{3}=x_{3}{ }^{\prime} & z=z^{\prime} \\
x_{4}=\gamma\left(i \beta x_{1}{ }^{\prime}+x_{4}{ }^{\prime}\right) & t=\gamma\left(\frac{\beta}{c} x^{\prime}+t^{\prime}\right)
\end{array}
$$

## 19. Doppler shift and aberration

19.1.



$$
k_{\mu}=\frac{p}{\hbar}=\left(k, \frac{i \omega}{c}\right)
$$

where $\omega=c k$

$$
k_{\mu} x_{\mu}=\mathbf{k} \cdot \mathbf{r}-\omega t=\text { invariant under the Lorentz transformation. }
$$

or

$$
k_{\mu} x_{\mu}=k x \cos \theta+k y \sin \theta-\omega t
$$

This should be equal to

$$
k_{\mu}^{\prime} x_{\mu}^{\prime}=k^{\prime} x^{\prime} \cos \theta^{\prime}+k^{\prime} y^{\prime} \sin \theta^{\prime}-\omega^{\prime} t^{\prime}
$$

Note that

$$
\begin{aligned}
& x=\gamma\left(x^{\prime}+v t^{\prime}\right) \\
& t=\gamma\left(t^{\prime}+\frac{v}{c^{2}} x^{\prime}\right) \\
& \omega=c k \\
& \omega^{\prime}=c k^{\prime}
\end{aligned}
$$

Substituting these parameters into the invariant form, we have

$$
t^{\prime}\left(c k^{\prime}-c k \gamma+k v \gamma \cos \theta\right)+x^{\prime}\left(-\frac{k v \gamma}{c}+k \gamma \cos \theta-k^{\prime} \cos \theta^{\prime}\right)+y^{\prime}\left(k \sin \theta-k^{\prime} \sin \theta^{\prime}\right)=0
$$

This should be satisfied for any $t^{\prime}, x^{\prime}$, and $y^{\prime}$.

$$
\begin{aligned}
& k \gamma(c-v \cos \theta)=c k^{\prime} \\
& k \gamma\left(\cos \theta-\frac{v}{c}\right)=k^{\prime} \cos \theta^{\prime} \\
& k \sin \theta=k^{\prime} \sin \theta^{\prime}
\end{aligned}
$$

### 19.2. Doppler shift

Since $k=\frac{2 \pi}{\lambda}, k^{\prime}=\frac{2 \pi}{\lambda^{\prime}}$

$$
\frac{2 \pi}{\lambda} \gamma\left(1-\frac{v}{c} \cos \theta\right)=\frac{2 \pi}{\lambda^{\prime}}
$$

or

$$
\lambda^{\prime}=\frac{\lambda}{\gamma\left(1-\frac{v}{c} \cos \theta\right)}
$$

### 19.3 Derivation of the formula using Mathematica

((Mathematica-14))

```
eq1=(k (x Cos[0]+y Sin[0])- \omega t -k1 (x1 Cos[01]+y1
Sin[01])+\omega1 t1)/.{\omega->c k,\omega1->c kl}//Simplify
    -c k t+c k1 t1+k x Cos[Ө]-k1 x1 Cos[01]+k y Sin[0]-k1
y1 Sin[01]
    rule1 = {x->\gamma(x1 + v t1), t > \gamma (t1 + \frac{v}{c}}\mathbf{c
    {x->(t1v + x1) r,t }->(t1+\frac{vx1}{\mp@subsup{c}{}{2}})\gamma,y->y1
    eq2=eq1/.rule1//Simplify
    ck1t1-ck(t1 + vx1
    k1 x1 Cos[01] +ky1 Sin[0] - k1y1 Sin[01]
    Collect[eq2,{x1,y1,t1}]
    t1 (ck1-ck\gamma + kv\gamma Cos[0]) +
    x1 (-\frac{kv\gamma}{c}+k\gamma\operatorname{Cos[0] - k1 Cos[01]})+y1(k\operatorname{Sin}[0]-k1\operatorname{Sin}[01])
```

Longitudinal Doppler shift

$$
\begin{aligned}
& \lambda 1[0] \\
& \frac{c \sqrt{1-\frac{v^{2}}{c^{2}}} \lambda}{c-v}
\end{aligned}
$$

$$
\text { Series }[\lambda 1[0],\{v, 0,4\}]
$$

$$
\lambda+\frac{\lambda v}{c}+\frac{\lambda v^{2}}{2 c^{2}}+\frac{\lambda v^{3}}{2 c^{3}}+\frac{3 \lambda v^{4}}{8 c^{4}}+0[v]^{5}
$$

Transverse Doppler shift

$$
\begin{aligned}
& \lambda 1\left[\frac{\pi}{\mathbf{2}}\right] \\
& \sqrt{1-\frac{\mathrm{v}^{2}}{\mathrm{c}^{2}}} \lambda
\end{aligned}
$$

$$
\operatorname{Series}\left[\lambda 1\left[\frac{\pi}{2}\right],\{v, 0,4\}\right]
$$

$$
\lambda-\frac{\lambda v^{2}}{2 c^{2}}-\frac{\lambda v^{4}}{8 c^{4}}+0[v]^{5}
$$

$$
\mathrm{f} 1=\frac{\lambda 1[\theta]}{\lambda} / .\{\mathbf{V} \rightarrow \mathbf{c} \beta\} / . \theta \rightarrow \frac{\pi \alpha}{180} / / \text { FullSimplify }
$$

$$
\frac{\sqrt{1-\beta^{2}}}{1-\beta \operatorname{Cos}\left[\frac{\pi \alpha}{180}\right]}
$$

Plot [Evaluate[Table[f1, $\{\beta, 0,0.99,0.05\}]$ ], $\{\alpha, 0,90\}$, Plots tyle $\rightarrow$ Table [Hue [0.05
i], $\{i, 0,20\}]$, Prolog $\rightarrow$ AbsoluteThickness [2], PlotRange $\rightarrow\{\{0$, 90\} 1 " $\}$ ] $\{0,7\}\}$, Background $\rightarrow$ GrayLevel $[0.7]$, AxesLabel $\rightarrow\{" \theta ", " \lambda$

$$
\begin{aligned}
& \text { rule } 2=\left\{k \rightarrow \frac{2 \pi}{\lambda}, k 1 \rightarrow \frac{2 \pi}{\lambda 1}\right\} \\
& \left\{k \rightarrow \frac{2 \pi}{\lambda}, k 1 \rightarrow \frac{2 \pi}{\lambda 1}\right\} \\
& \text { eq3=(c kl-c k } \gamma+\mathrm{k} \text { v } \gamma \text { Cos [Ө])/.rule2//Simplify } \\
& \frac{2 \pi\left(C(\lambda-\gamma \lambda 1)+v_{\gamma} \lambda 1 \operatorname{Cos}[\theta]\right)}{\lambda \lambda 1} \\
& e q 4=-\frac{k v \gamma}{c}+k \gamma \operatorname{Cos}[\theta]-k 1 \operatorname{Cos}[\theta 1] / . \text { rule2 // Simplify } \\
& -\frac{2 \pi\left(V_{\gamma} \lambda 1-\mathbf{C} \gamma \lambda 1 \operatorname{Cos}[\theta]+\mathrm{c} \lambda \operatorname{Cos}[\theta 1]\right)}{\mathrm{c} \lambda \lambda 1} \\
& \text { eq5 } 5 \text { k } \operatorname{Sin}[\theta]-k 1 \text { Sin[Ө1]/.rule2//Simplify } \\
& \frac{2 \pi \operatorname{Sin}[\theta]}{\lambda}-\frac{2 \pi \operatorname{Sin}[\theta 1]}{\lambda 1} \\
& \text { eq31 =Solve [eq3 }=0, \lambda 1] / / \text { Simplify } \\
& \left\{\left\{\lambda 1 \rightarrow \frac{\mathrm{c} \lambda}{\mathrm{c} \gamma-\mathrm{v} \gamma \operatorname{Cos}[\theta]}\right\}\right\} \\
& \lambda 1\left[\theta_{-}\right]=\lambda 1 / . \operatorname{eq} 31[[1]] / .\left\{\gamma \rightarrow \frac{1}{\sqrt{1-\frac{v^{2}}{c^{2}}}}\right\} / / \text { Simplify } \\
& \frac{c \sqrt{1-\frac{v^{2}}{c^{2}}} \lambda}{c-v \operatorname{Cos}[\theta]}
\end{aligned}
$$



## 19.4. longitudinal Doppler shift $\theta=0$ (red shift)

We suppose that a source is located at the origin of the reference frame $S$. An observer moves relative to $S$ at velocity $v$. So that he is at rest in $S^{\prime}$.
$\lambda^{\prime}=\frac{\lambda}{\gamma\left(1-\frac{v}{c}\right)}=\lambda \sqrt{\frac{1+\frac{v}{c}}{1-\frac{v}{c}}}$
If $S$ ' moves toward S , rather than away from S , the signs in numerator and denominator of the radical would have been unterchanged.
((The red shift)) Wikipedia
The light from distant stars and more distant galaxies is not featureless, but has distinct spectral features characteristic of the atoms in the gases around the stars. When these spectra are examined, they are found to be shifted toward the red end of the spectrum. This shift is apparently a Doppler shift and indicates that essentially all of the galaxies are moving away from us. Using the results from the nearer ones, it becomes evident that the more distant galaxies are moving away from us faster. This is the kind of result one would expect for an expanding universe.

The building up of methods for measuring distance to stars and galaxies led Hubble to the fact that the red shift (recession speed) is proportional to distance. If this proportionality (called Hubble's Law) holds true, it can be used as a distance measuring tool itself.

The measured red shifts are usually stated in terms of a z parameter. The largest measured z values are associated with the quasars.
((Mathematica-16))
Red shift: $\lambda^{\prime} / \lambda$ vs $\beta=\mathrm{v} / \mathrm{c}$

$$
\begin{aligned}
& \text { eq } 1=\sqrt{\frac{1+\beta}{1-\beta}} \\
& \sqrt{\frac{1+\beta}{1-\beta}}
\end{aligned}
$$

```
    Plot[eq1,{ , 0.4,0.99},
PlotStyle }->{\mathrm{ Hue [0],Thickness[0.015]},
Background->GrayLevel[0.7],AxesLabel->{"\beta","\lambda'/\lambda"}]
```



```
    -Graphics-
z-parameter = ( }\mp@subsup{\lambda}{}{\prime}-\lambda)/\lambda=(\mp@subsup{\lambda}{}{\prime}/\lambda)-
    z=eq1-1
    -1+\sqrt{}{\frac{1+\beta}{1-\beta}}
    Plot [z,{ \beta,0.4,0.99},
PlotStyle }->{\mathrm{ Hue [0.5],Thickness[0.015]},
Background->GrayLevel[0.7],AxesLabel }->{|\beta","z"}
*)
```

19.5. Transverse Doppler shift ( $\theta=\frac{\pi}{2}$ )

$$
\lambda^{\prime}=\frac{\lambda}{\gamma}=\lambda \sqrt{1-\frac{v^{2}}{c^{2}}}
$$

### 19.6. Aberration

From these equations

$$
\begin{aligned}
& k \gamma\left(1-\frac{v}{c} \cos \theta\right)=k^{\prime} \\
& k \gamma\left(\cos \theta-\frac{v}{c}\right)=k^{\prime} \cos \theta^{\prime}
\end{aligned}
$$

we have

$$
\cos \theta^{\prime}=\frac{\cos \theta-\frac{v}{c}}{1-\frac{v}{c} \cos \theta}
$$

For low velocity we can neglect $v^{2} / c^{2}$ and higher-order terms. Setting $\theta^{\prime}=\theta+\Delta \theta$

$$
\cos (\theta+\Delta \theta)=\cos \theta-\Delta \theta \sin \theta
$$

and

$$
\frac{\cos \theta-\frac{v}{c}}{1-\frac{v}{c} \cos \theta}=\cos \theta-\beta \sin ^{2} \theta
$$

Then we have

$$
\Delta \theta=\beta \sin \theta
$$

((Mathematica-17))

```
Aberration
eq1 \(=\operatorname{Cos}[\theta+\Delta \theta]\)
\(\operatorname{Cos}[\Delta \theta+\theta]\)
Series [eq1, \(\{\Delta \theta, 0,1\}]\)
\(\operatorname{Cos}[\theta]-\operatorname{Sin}[\theta] \Delta \theta+0[\Delta \theta]^{2}\)
eq11= \(\operatorname{Cos}[\theta]-\operatorname{Sin}[\theta] \Delta \theta\)
\(\operatorname{Cos}[\theta]-\Delta \theta \operatorname{Sin}[\theta]\)
eq \(2=\frac{\operatorname{Cos}[\theta]-\beta}{1-\beta \operatorname{Cos}[\theta]}\)
\(\frac{-\beta+\operatorname{Cos}[\theta]}{1-\beta \cos [\theta]}\)
Series [eq2, \(\{\beta, 0,1\}\) ]
\(\cos [\theta]+\left(-1+\operatorname{Cos}[\theta]^{2}\right) \beta+0[\beta]^{2}\)
eq22 \(=\operatorname{Cos}[\theta]+\left(-1+\operatorname{Cos}[\theta]^{2}\right) \beta\)
\(\operatorname{Cos}[\theta]+\beta\left(-1+\operatorname{Cos}[\theta]^{2}\right)\)
eq3 \(=\) eq11 \(=\) eq2 2
\(\operatorname{Cos}[\theta]-\Delta \theta \operatorname{Sin}[\theta]=\operatorname{Cos}[\theta]+\beta\left(-1+\operatorname{Cos}[\theta]^{2}\right)\)
eq31=Solve [eq3, \(\Delta \Theta\) ] //Simplify
```

```
General ::spell1 : Possible spelling error : new
    symbol name "eq31" is similar to existing symbol "eq1". More..
{{\Delta0->\beta Sin [0] }}
```


## REFERENCES

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## Appendix

## A-1 Capacitance

The capacitor moves at the constant speed in the $x$ direction.
In the S' system

$$
\begin{aligned}
& E_{2}^{\prime}=\frac{\sigma_{0}}{\varepsilon_{0}} \\
& E_{3}^{\prime}=0 \\
& E_{1}^{\prime}=0
\end{aligned}
$$

From the Lorentz transformation, the electric field and the magnetic field in the S system are given by

$$
\begin{aligned}
& E_{1}=E_{1}^{\prime}=0 \\
& E_{2}=\gamma\left(E_{2}{ }^{\prime}+c \beta B_{3}{ }^{\prime}\right)=\gamma E_{2}^{\prime}=\gamma \frac{\sigma_{0}}{\varepsilon_{0}} \\
& E_{3}=\gamma\left(E_{3}{ }^{\prime}-c \beta B_{2}{ }^{\prime}\right)=0
\end{aligned}
$$

$$
\begin{aligned}
& B_{1}=B_{1}{ }^{\prime}=0 \\
& B_{2}=\gamma\left(B_{2}{ }^{\prime}-\frac{\beta}{c} E_{3}{ }^{\prime}\right)=0 \\
& B_{3}=\gamma\left(B_{3}{ }^{\prime}+\frac{\beta}{c} E_{2}{ }^{\prime}\right)=\gamma \frac{\beta}{c} E_{2}{ }^{\prime}=\gamma \frac{\beta}{c} \frac{\sigma_{0}}{\varepsilon_{0}}
\end{aligned}
$$

Since $E_{2}$ is expressed by

$$
E_{2}=\frac{\sigma}{\varepsilon_{0}}=\gamma \frac{\sigma_{0}}{\varepsilon_{0}}
$$

or
$\sigma=\gamma \sigma_{0}$

## A-2 Faraday's law

## 1. A conducting rod moving through a uniform magnetic field

We consider a metal rod (conductor) which moves at a constant velocity ( $\boldsymbol{v}$ ) in a direction perpendicular to its length. Pervading the space through which the rod moves there is a uniform magnetic field $\boldsymbol{B}(/ / z)$ constant in time. There is no electric field in the reference frame F .

The rod contains charged particles that will move if a force is applied to them. Any charged that is carried along with the rod, such as the particle of charge $q$ moves through the magnetic field $\mathbf{B}$ and thus experience a force.

$$
\mathbf{f}=q(\mathbf{v} \times \mathbf{B})
$$

The direction of the force is dependent on the sign of the charge $q$.
When the rod is moving at constant speed and things have settled to a steady state, the force $\boldsymbol{f}$ must be balanced, at every point inside the rod, by an equal and opposite force. This can only arise from an electric field in the rod. The electric field develops in the following way. The force $\boldsymbol{f}$ pushes negative charges toward one end of the rod, leaving the other end positively charges. This goes on until these separated charges themselves cause an electric field $\boldsymbol{E}$ such that, everywhere in the interior of the rod,

$$
q \mathbf{E}+\mathbf{f}=0,
$$

Then the motion of charge relative to the rod ceases. This charge distribution causes an electric field outside the rod, as well as inside. Inside the rod, there has developed an
electric field $\mathbf{E}=\mathbf{v} \times \mathbf{B}_{z}$, exerting a force $q \boldsymbol{E}$ which just balances the force

$q(\mathbf{v} \times \mathbf{B}) . \times$
Let us observe the system from a frame F' that moves with the rod. What is the magnetic field $\mathbf{B}$ ' and the electric field $\mathbf{E}$ '? Note that there is no electric field $(\boldsymbol{E}=0)$ in the frame $F(\mathbf{B} \neq 0)$. The $\boldsymbol{E}^{\prime}$ and $\boldsymbol{B}^{\prime}$ in the frame $F^{\prime}$ are related to those in the frame $F$ as

$$
\begin{array}{ll}
E_{1}^{\prime}=E_{1}=0 & B_{1}{ }^{\prime}=B_{1}=0 \\
E_{2}^{\prime}=\gamma\left(E_{2}-c \beta B_{3}\right)=-\gamma B_{3}, & B_{2}{ }^{\prime}=\gamma\left(B_{2}+\frac{\beta}{c} E_{3}\right)=\gamma B_{2}=0 \\
E_{3}^{\prime}=\gamma\left(E_{3}+c \beta B_{2}\right)=\gamma B_{2}=0 & B_{3}{ }^{\prime}=\gamma\left(B_{3}-\frac{\beta}{c} E_{2}\right)=\gamma B_{3}
\end{array}
$$

or

$$
\mathbf{E}^{\prime}=\gamma(\mathbf{v} \times \mathbf{B})=\mathbf{v} \times \mathbf{B}^{\prime}
$$

where $\gamma=\frac{1}{\sqrt{1-v^{2} / c^{2}}} \approx 1$ for $v \ll c$. The magnetic field $\mathbf{B}^{\prime}(=\gamma \boldsymbol{B})$ is almost equal to $\boldsymbol{B}$. The electric field $\boldsymbol{E}^{\prime}$ has only a component along the $y^{\prime}$ axis (the same as $y$ xis). The presence of the magnetic field $\boldsymbol{B}^{\prime}$ has no influence on the static charge distribution.


## 2. A loop moving through a nonuniform magnetic field


$\boldsymbol{F}$ denotes the force which acts on a charge q that rides along with the loop. We evaluate the line integral of $\boldsymbol{F}$, taken around the whole loop. On the two sides of the loop which lie parallel to the direction of motion, $\boldsymbol{F}$ is perpendicular to the path element ds. So there is no contribution to the line integral. Taking account of the contributions from the other two sides, each of length $w$, we have

$$
W_{\text {net }}=\oint \mathbf{F} \cdot d \mathbf{s}=\oint q(\mathbf{v} \times \mathbf{B}) \cdot d \mathbf{s}=q v\left(B_{a}-B_{b}\right) w
$$

where

$$
\begin{aligned}
& \mathbf{F}_{a}=q\left(\mathbf{v} \times \mathbf{B}_{a}\right) \\
& \mathbf{F}_{b}=q\left(\mathbf{v} \times \mathbf{B}_{b}\right)
\end{aligned}
$$

From the definition, we get

$$
W_{\text {net }}=q V=q \oint \mathbf{E} \cdot d \mathbf{s}=q v\left(B_{a}-B_{b}\right) w
$$

or

$$
V=\oint \mathbf{E} \cdot d \mathbf{s}=v\left(B_{a}-B_{b}\right) w
$$

The electromotive force $V$ is related in a very simple way to the rate of change of magnetic flux through the loop. The magnetic flux through a loop is the surface integral of $\boldsymbol{B}$ over a surface which has the loop for its boundary.

$$
\Delta \Phi=B_{b} v w \Delta t-B_{a} v w \Delta t=-\left(B_{a}-B_{b}\right) v w \Delta t
$$

or

$$
\begin{aligned}
V & =\oint \mathbf{E} \cdot d \mathbf{s}=\oint(\nabla \times \mathbf{E}) \cdot d \mathbf{a}=-\frac{d \Phi}{d t}=-\frac{d}{d t} \int \mathbf{B} \cdot d \mathbf{a} \\
& =\left(B_{a}-B_{b}\right) v w
\end{aligned}
$$

We now consider the frame F' attached to the loop.


From the Lorentz transformation of $\boldsymbol{E}$ and $\boldsymbol{B}$,

$$
\begin{array}{ll}
E_{1}{ }^{\prime}=E_{1}=0 & B_{1}{ }^{\prime}=B_{1}=0 \\
E_{2}{ }^{\prime}=\gamma\left(E_{2}-v B_{3}\right)=-\gamma B_{3}, & B_{2}{ }^{\prime}=\gamma\left(B_{2}+\frac{\beta}{c} E_{3}\right)=\gamma B_{2}=0 \\
E_{3}{ }^{\prime}=\gamma\left(E_{3}+v B_{2}\right)=\gamma v B_{2}=0 & B_{3}{ }^{\prime}=\gamma\left(B_{3}-\frac{\beta}{C} E_{2}\right)=\gamma B_{3}
\end{array}
$$

we have

$$
\begin{array}{ll}
E_{a}{ }^{\prime}=-\gamma B_{a} \approx-v B_{a}, & E_{b}{ }^{\prime}=-\gamma B_{b}=-v B_{b} \\
B_{a}{ }^{\prime}=\gamma B_{a} \approx B_{a} & B_{b}{ }^{\prime}=\gamma B_{b} \approx B_{b}
\end{array}
$$

For observers in the frame $S^{\prime}, \boldsymbol{E}_{\mathrm{a}}{ }^{\prime}$ and $\boldsymbol{E}_{\mathrm{b}}{ }^{\prime}$ are genuine electric field. It is not an electrostatic field. The integral of $\boldsymbol{E}^{\prime}$ around the loop, which is the electromotive force $V$, is given by

$$
V^{\prime}=^{\prime} \oint \mathbf{E} \cdot d \mathbf{s}=v B_{a} w-v B_{b} w=v w\left(B_{a}-B_{b}\right)
$$

which is the same as that obtained for the frame S .

